

1 **Exploratory Description of Low Frequency Ocean Temperature Variability**
2 **and Its Global Structure**

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ABSTRACT

8 An exploratory description is made of nearly global potential temperature
9 variability from months to 20 years using the field derived from the EC-
10 COv4 state estimate. Relative to higher frequencies, longer periods do not
11 exhibit simplification in the space-time structures. Frequency spectra at indi-
12 vidual points are usefully described by a reduced form of a previously pro-
13 posed analytic model. In contrast, the vertical structure of the variability at
14 low frequencies—periods beyond 1 year—has a very complex form, with only
15 a few global generalizations apparent. Meridional wavenumber spectra, re-
16 flecting the dominant zonality of oceanic low frequencies, are spatially com-
17 paratively simple, while the zonal wavenumber spectra are spatially complex
18 and not very meaningful. The emergence of strong spatial structures at longer
19 periods is consistent with the presence of complex time-mean (0 frequency)
20 structures in bottom topography, sidewalls, and meteorological forcing.

21 **1. Introduction**

22 This paper began as an attempt to realistically calculate the accuracies and precisions of pub-
23 lished estimates of global mean oceanic temperature (heat content) change through time. As
24 quickly became clear, the uncertainties of those calculations are dominated by the under-sampling
25 of space-time structures of low-frequency oceanic thermal variability. That in turn led to the need
26 to describe quantitatively the variability on a global scale at periods extending out to decades. No
27 such previously published description appears to exist. Thus what follows is a draft description
28 of low frequency oceanic thermal variability extending to 20-years duration. This effort raises
29 fundamental questions about how to describe (and then to understand and use) in a basic form, the
30 complex structure of a globally time and space varying fluid—if it is possible.

31 As will be seen, a large number of assumptions are necessary to proceed, and as a strawman, it
32 would be no surprise to find major future changes being made in the results. In the meantime, the
33 description does have exemplary uses both for the heat content problem and as tests of numerical
34 models directed at the ocean in climate. Application of the results to the estimates of heat content
35 change will be described elsewhere (Wunsch 2019).

36 As global data sets and global general circulation models with quantifiable skill have emerged
37 over the past few decades, the conclusion that the ocean has a very strong regional and temporal
38 complexity has become inescapable. Such technically beautiful ideas such as Sverdrup balance,
39 abyssal recipes, Stommel-Arons flows, et al., are increasingly perceived to have at best regional
40 applicability. Nonetheless a search for widely applicable principles describing oceanic physics
41 is worthwhile. So for example, Sonnewald et al. (2019), using a vertically integrated vorticity
42 balance, divided the ocean laterally into 6+ distinct dynamical regions, of greatly varying area.

43 Similarly, the altimetric wave number power law results of Xu and Fu (2012) seem to imply a
44 minimum of about 14 dynamical regimes.

45 Representations via spectral methods are both a useful summary description and have many
46 potential applications, depending upon the particular physical variable. The most familiar of
47 such spectra, in frequency, wavenumber, or both, are those describing surface gravity waves (e.g.,
48 Komen et al. 1994) and the internal wave spectrum in the various Garrett and Munk (GM) esti-
49 mates (Munk 1981 and many subsequent papers, e.g., Polzin and Lvov (2011)).

50 For the mesoscale, as documented in numerous publications, no universality comparable to that
51 seen for internal waves is found. In particular, the energy levels, as determined primarily from
52 altimetric data, vary by more than two orders of magnitude (e.g., Hughes et al. 2018; their Fig. 1).
53 A spectral representation has many useful applications, at least in a local sense. For Gaussian pro-
54 cesses stationary in both space and time, the spectral representation along with the various mean
55 values, is a complete description of the stochastic field. Here, the spatial inhomogeneity means that
56 the frequency-wavenumber spectrum is not, even for purely Gaussian data, a complete stochastic
57 representation—it is the first term in a higher order series. Nonetheless, its descriptive power, at
58 least locally, remains useful. As one example, Fig. 1 shows the histogram of monthly potential
59 temperature anomalies (relative to the 20-year mean) at one depth (477m) from an oceanic state
60 estimate (e.g., Fukumori et al. 2018) described below. Although no formal test of normality has
61 been made, the near- symmetric, unimodal distribution permits a ready interpretation of standard
62 deviations and variances.

63 The focus here is on the temperature spectrum, because it is a scalar, has a very long history of in
64 situ measurements, is important climatologically and is, in recent times in the upper ocean, widely
65 sampled by the Argo array. But direct observations of deep (below about 2000m) interannual
66 variability are very few—in a volume of roughly half the ocean. What follows relies on those data

67 that do exist, and on the physics connecting the upper and abyssal oceans. Results are considered
68 “exploratory” both because the full skill of the estimates at the longest periods is not known, and
69 the best way to describe and use the resulting complicated behavior of a time-varying three-space-
70 dimensional global field is also unclear.

71 **2. The State Values and Their Frequency Spectra**

72 Wortham and Wunsch (2014; hereafter WW14) proposed a general form of a three-space-
73 dimension and time-spectrum for low frequency (below the Coriolis frequency, f) oceanic vari-
74 ability and compared it to a variety of data, both altimetric and from in situ measurements.

75 Their form followed on earlier discussions of Zang and Wunsch (2001; hereafter ZW01). WW14
76 showed a quantitatively useful fit to a wide variety of data, including altimetry, and from moored
77 measurements of temperature and velocity. Subsequent developments for analogous statistical
78 descriptions include Wortham et al. (2014); Abernathey and Wortham (2015); Samelson et al.
79 (2016). Altimetry data have dominated the discussion of the wavenumber components of the
80 spectrum, although for high wavenumbers (wavelengths shorter than about 100-200km), consider-
81 able deviation from a universal form is known (e.g., Xu and Fu 2012, but see Callies and Wu 2019,
82 for discussion of their interpretation). Here the WW14 form is used to describe the temperature
83 temporal-frequency spectrum so as to draw some near-global conclusions.

84 The values used here are from the twenty-year time-varying subset climatology of the Estimating
85 the Circulation and Climate of the Ocean version 4 (ECCOv4) state estimate (see e.g., Fukumori
86 et al. 2018 and references there). This climatology represents a weighted least-squares fit of the
87 MITgcm (Marshall et al. 1997) and its evolutionary ECCO successors, to the diverse near-global
88 data sets and meteorological forcing estimates that became available during and after the World
89 Ocean Circulation Experiment. The single most important feature of this model representation is

90 that the model is *free* running but with its numerous control parameters having been previously
91 adjusted so that the model trajectory takes it through all of the data points within (mostly) estimates
92 of their uncertainties.

93 As with many global analyses, the chief obstacle here is that the model is non-eddy-resolving
94 (with a 1 degree of longitude and a variable latitude grid). Thus the strong assumption is, nonethe-
95 less, made here that the solution fit on periods exceeding several months and wavelengths ex-
96 ceeding $\approx 200\text{km}$ provides a quantitative estimate of the variability, its nature and structure. That
97 assumption in turn rests upon the supposition that the dominant quasi-geostrophic nature of the
98 ocean circulation is well-captured by the data—and hence the adjusted model—at long periods—
99 an inference consistent with the comparisons to moored data in WW14. In particular, those authors
100 showed that frequency spectra within the eddy band appeared to be smooth extensions of the spec-
101 tra at lower frequencies. Hence one of the assumptions made here is that the low frequency results
102 can be extrapolated into the eddy band.

103 To the extent that low frequency features produced e.g., by eddy-eddy or eddy-mean flow in-
104 teractions, remain in thermal wind balance, the state estimate will properly represent them. The
105 assumption is equivalent to the assertion that away from boundaries,

$$\frac{\partial}{\partial t} \left(f \frac{\partial v}{\partial z} - g \frac{\partial \rho}{\partial x} \right) = O(\varepsilon) \quad (1)$$

106 where ε is a small number relative to the left-hand-side terms and the fields are varying over
107 years and longer. Boundary layers are not properly resolved with the existing resolution and
108 a further assumption is that the unresolved regions are passively consistent with the strongly
109 constrained, quasi-geostrophic, interior circulation and structure (an assumption underlying the
110 Stommel-Arons picture). Penduff et al. (2010) and others show the integrated effects of eddies
111 on the large-scale circulation; the extent to which a geostrophic pressure field is induced to bal-

112 ance that circulation is not clear. Balanced eddy structures, particularly those associated with
113 topographic features, can persist for periods much exceeding a year (no low-frequency cut-off is
114 known), and should future studies show major ageostrophic physics in eddy-resolving runs at low
115 frequencies, what follows would have to be re-evaluated.

116 *Frequency Spectra of the State Estimate*

117 From the monthly average values of ECCOv4 at 477m the spectral density ranging from 20 years
118 to 2 months (the Nyquist period) is readily computed as a function of frequency ω' . A depth of
119 477m was chosen as a reasonable global compromise value lying primarily below the mixed layer
120 and above the main thermocline. The standard deviation of monthly anomalies averaged over 20
121 years at that depth is shown in Fig. 1. Converted to power, spectra sum to the squares of these
122 values.

123 The discussion that follows is restricted to the region northward of 40°S , as the Southern Ocean
124 with its strong mean advecting eastward flow is spectrally distinct from the remaining oceans
125 (WW14). To the north, the sea ice region poleward of about 55°N is also omitted. This restriction
126 still leaves many special dynamical regions in the domain, including the tropics. Conventional
127 spectral estimates using a Daniell (rectangular frequency or wavenumber) window were computed
128 for 27,634 distinct locations where the depth was at least 500m. An estimated $\nu = 6$ degrees of
129 freedom was used at each spectral estimate frequency. The median (not area weighted) of all of
130 these spectra can be seen in Fig. 2. In many locations, and as appears in the median result, the
131 annual cycle and sometimes its harmonics, is conspicuous. In the net power and fitting results,
132 these peaks are ignored because fitting without them changed the results only slightly. Its pres-
133 ence may be compensatory for a possible underestimate of power between two months and a year.
134 The median power at 477m is $0.10(^{\circ}\text{C})^2 = (0.31\text{ }^{\circ}\text{C})^2$ from 20 years to two months. A rough

135 description would be that it has a ω'^{-2} behavior at frequencies above the annual cycle, and is flat-
 136 tened, tending toward white noise, at lower frequencies. The whitish character is consistent with
 137 the absence of strong (relative to the variability) data trends. A separate study of the temperature
 138 annual cycle and its overtones would be worthwhile.

139 *The Analytic Spectrum*

140 The WW14 spectrum is an empirical one for the interior ocean only, whose construction was
 141 guided only by general discussions both in theory and observation of wavenumber and frequency
 142 behavior (e.g., Vallis 2017; Arbic et al. 2012; Scott et al. 2017). In local Cartesian horizontal
 143 coordinates and time, it is of the form,

$$\Phi_{\psi}(k', l', \omega', x, y) = \tag{2}$$

$$A \left\{ \frac{1}{(k'^2 L_x^2 + l'^2 L_y^2 + 1)^\alpha (\omega'^2 T^2 + 1)} + \exp(-[k'^2 L_x^2 + l'^2 L_y^2 + T^2(k' c_x + l' c_y - \omega')^2]) \right\}.$$

144 Here, k', l', ω' are *non-radian* wavenumbers and frequencies, L_x, L_y, T are the zonal and meridional
 145 spatial scales and a temporal scale. c_x, c_y are empirical phase speeds, $\alpha \approx 5/2$. A is an adjustable
 146 overall magnitude. As discussed by WW14, all these parameters are functions of position—so
 147 that using local Cartesian coordinates makes some sense. Dependence upon x, y is slowly varying
 148 by assumption and the coordinates are usually suppressed below, but are implicit. The subscript
 149 ψ denotes the stream function as defined by ZW01. Each physical variable, be it temperature,
 150 salinity, surface elevation, velocity has a differing power density spectrum where a multiplying
 151 factor converts Φ_{ψ} into the applicable form. Thus the temperature spectrum would be,

$$\Phi_{\theta}(k', l', \omega', z, m) = \left(f \frac{\partial \langle \theta_0 \rangle(z, x, y)}{\partial z} G_m(z, x, y) \right)^2 \Phi_{\psi}(k', l', \omega'). \tag{3}$$

152 Here, $G_m(z, x, y)$ is the m -th vertical displacement mode. $\partial \langle \theta_0 \rangle / \partial z$ is the time average local ver-
 153 tical temperature derivative. Eqs. (2, 3) are over-simplified compared to reality. On the other hand,

154 WW14 show that subject to regional adjustment of the various parameters, they are quantitatively
 155 useful for a variety of time and space-time series.

156 The first term on the right in Eq. (2) represents the frequency-wave-number continuum including
 157 equal amounts of energy moving both eastward and westward, and northward and southward and
 158 is similar to the ZW01 form. The second term on the right represents the asymmetric westward-
 159 going energy dominated at low frequencies by the so-called non-dispersive line (NDL) whose
 160 slope in $k' - \omega'$ space is controlled by c_x (e.g., WW14). The NDL is conspicuous in altimetric
 161 data, in large-part because of its strong barotropic component and by its non-linear coupling to a
 162 strongly surface-amplified first baroclinic mode velocity. It is far less prominent in temperatures
 163 measured or computed at depth. From the altimetric data, its structure is imposed upon the state
 164 estimate at low frequencies and wavenumbers. Whether it is more wave-like or more isolated
 165 vortex-like (Chelton et al. 2011) over all frequencies and wavenumbers has not been explored.
 166 In either case, when viewed through a low-pass frequency and wavenumber filter, the longest
 167 wavelengths and periods are seen. Note that from frequency spectra alone at individual points,
 168 one cannot distinguish propagating from standing energy.

169 When Φ_ψ is integrated over all k', l' , the frequency spectrum is obtained (WW14; their Eq. 33):

$$\Omega_\theta(\omega') = A \left\{ \frac{\pi}{(\alpha - 1)L_x L_y} \frac{1}{(1 + \omega'^2 T^2)} + \frac{\pi}{\sqrt{D}} \exp(-L_x^2 L_y^2 T^2 \omega'^2 / D) \right\}, \quad (4)$$

$$D = c_x^2 L_y^2 T^2 + L_x^2 L_y^2 + c_y^2 L_x^2 T^2$$

170 where the second term represents the frequency spectrum of the NDL. Here A absorbs the factor
 171 $(f \partial \langle \theta_0 \rangle / \partial z)^2$. In fitting this frequency form, globally, two changes are made: combined param-
 172 eters are lumped, and the $1 / (1 + T^2 \omega'^2)$ term and the NDL term are separated to provide an extra
 173 degree of freedom in accounting the known latitudinal composition of Rossby waves. Thus the fit

174 is to:

$$\Omega'_\theta(\omega') = \frac{A_1}{(1 + \omega'^2 T^2)} + A_2 \exp(-\omega'^2/P_a) \quad (5)$$

175 where locally,

$$A_1 = \frac{A\pi}{(\alpha - 1)L_x L_y}, A_2 = \frac{A\pi}{\sqrt{D}}, P_a = \frac{D}{L_x^2 L_y^2 T^2}, \quad (6)$$

176 and where the fitting parameters are A_1, T, A_2, P_a . The first term represents a process obeying an
177 ω'^{-2} power law at high frequencies and is intimately related to that of a continuous time AR(1)
178 process (e.g., Hughes and Williams, 2010). Partial derivatives of Eq. (5) show that the value of
179 A_1 depends upon T , etc. The second, exponential, term from the NDL, decays faster in frequency
180 than does any algebraic power. Spectra that are power laws of ω'^{-2} and steeper in some range
181 $\omega'_1 \leq \omega' \leq \infty$ are insensitive to aliasing from frequencies above ω'_1 (see Wunsch 1972, Eq. 13;
182 Rhines and Huybers 2011).

183 A fit was made of $\log \Omega'_\theta(\omega', x, y)$ to the logarithm of the empirical temperature spectra of
184 ECCOV4 over the range of periods from 20 years to two months. Logarithms are used to render
185 the spectral values more nearly constant with frequency.¹ The resulting parameters vary widely
186 over the global ocean as seen in Figs. 3 - 4.

187 Overall misfits are generally within about 6% of the state estimate log spectrum relative to the
188 log of the fit spectrum, but not always. Notable features as might have been expected are: (1) Val-
189 ues are spatially noisy, consistent with a comparatively short 20-year time-interval; (2) amplitudes
190 A_1 are small in the tropical oceans with exceptions on eastern boundaries of the Atlantic and In-
191 dian Oceans; (3) High latitudes show enhancement of A_1 . (4) Over most of the ocean T has values
192 of a few years, with much longer values at high latitudes, and much lower values at low latitudes.
193 More generally, the regions of high values of T are where the spectra tend to be red out to the

¹Fitting was through a nonlinear trust region method (Seber and Wild 1989, P. 603+; Mathworks Website 2018), a form of least-squares using a search over a variable neighborhood about the most recent estimate.

194 longest periods. Whether these represent true rednoise processes or the presence of trends has to
195 be examined separately. (5) The contribution of the NDL to temperature is slight almost every-
196 where. (Exceptions exist to all of these sweeping summaries.) Because of the equivalent spatial
197 complexity and the much smaller amplitudes, the fits to the NDL line are shown in the Appendix.
198 Median values from the fits are $A_1 = 0.070^\circ\text{C}^2/\text{cycle}/\text{year}$, $T = 10.3\text{y}$, $A_2 = 0.0048^\circ\text{C}^2/\text{cycle}/\text{year}$,
199 $P_a = 1.3\text{y}^2$, differing somewhat from the direct fit to the median spectrum (a result expected from
200 a nonlinear estimation process).

201 An approximate uncertainty of the fitting result can be determined from the inverse Hessian
202 based upon its local value calculated from the Jacobian, which is a by-product of the optimization
203 algorithm (Kalmikov and Heimbach 2014). When multiplied by the covariance matrix of the data
204 noise, it provides an estimate of the uncertainty. The main result is that the 4 parameter estimates
205 have strong and highly variable correlated uncertainties. A global display of the individual 4x4
206 matrices at each grid point is not easy to digest and is omitted here.

207 **3. Temporal Autocovariance**

208 Oceanic physical processes have a strong frequency dependence, and that leads to the central
209 importance of spectral analyses. On the other hand, for calculating e.g., the expected accuracy of a
210 space or time or space-time average or map, the integrated time-scales included in the covariances
211 are a more convenient tool.

212 By the Wiener-Khinchin theorem, the temporal autocovariance, $R(\tau)$, and the autocorrelation,
 213 $\phi(\tau)$, of the frequency spectrum Eq. (5) are,

$$\begin{aligned}
 R(\tau) &= \frac{1}{2\pi} \int_{-\infty}^{\infty} \Omega'_{\theta}(\omega') \cos(\omega' \tau) d\omega' & (7a) \\
 &= \frac{1}{\pi} \int_0^{\infty} \left[\frac{A_1}{(1 + \omega'^2 T^2)} + A_2 \exp(-\omega'^2 / P_a) \right] \cos(2\pi \omega' \tau) d\omega' \\
 \phi(\tau) &= R(\tau) / R(0) \\
 &= \frac{A_1 \pi / (2T) \exp(-\tau / T) + A_2 P_a^{1/2} \pi^{1/2} \exp(-P_a \tau^2 / 4)}{A_1 \pi / (2T) + A_2 P_a^{1/2} \pi^{1/2}}, \quad \tau \geq 0
 \end{aligned}$$

214 Decay time to effective zero correlation ($\phi(\tau_d) = 0.1$) is shown in Fig. 5 and which can be
 215 interpreted as providing the temporal separation required for statistically independent temperature
 216 samples at this depth. Tropical areas have times of order one year, while patches (including the
 217 central equatorial Pacific) take more than 10 years to decorrelate. This decorrelation time is im-
 218 portant in calculations of the accuracy of large-scale sample averages. In low latitudes, samples
 219 obtained two years apart could be deemed independent, whereas at high latitudes that can take 10
 220 or more years. The median value is 2.8 years and the mean 3.6 years, omitting values where more
 221 than 20 years is required.

222 4. Wavenumber Spectra

223 Computation of wavenumber spectral densities involves choosing distances (or areas) over
 224 which they are representative, and the presence of complex land boundaries does not lend itself
 225 to easy or automatic selection. Here what is done is to separately determine the zonal and merid-
 226 ional wavenumber periodograms along each line of latitude or longitude extending 30° eastward
 227 or northward. The calculations are done for each year separately, and then averaged over 20 years
 228 to give a power spectrum estimate. Nominal position is assigned to the mid-point longitude or
 229 latitude.

230 Again simplifying the WW14 forms (their Eq. 31) , now omitting the contribution of the NDL
 231 which is a pure exponential, the results are then fit to the continuum,

$$\Phi(\mathbf{r}_j, k') = B(1 + L_y^2 k'^2)^{1/2 - \alpha} \quad (8)$$

232 for the zonal spectra, and to

$$\Phi(\mathbf{r}_j, l^{prime}) = C(1 + L_x^2 l'^2)^{1/2 - \alpha} \quad (9)$$

233 for the meridional. Note the pairing of k' with L_y and of l' with L_x , consistent with quasi-
 234 geostrophic balance. At high wavenumbers, the behavior is a power law with exponent $1 - 2\alpha < 0$,
 235 if $\alpha > 1/2$, again becoming more white at long wavelengths.

236 *Meridional Wave Numbers*

237 The meridional wavenumber spectrum reflects, in the thermal wind/geostrophic balance, the
 238 zonal flows. A large literature exists describing and rationalizing a tendency for the circulation to
 239 have preferred zonality, and in the extreme of the appearance of zonal jets (e.g., Chen et al. 2015;
 240 Galperin and Reid 2019) . Results of fitting Eq. (9) are displayed in Figs. 6-7. Apart from the
 241 prominent western intensification visible in C , both L_x and α are remarkably uniform and stable
 242 (α is displayed in the Appendix). Both the mean and median value of $\alpha \approx 3/2$ in contrast to the
 243 estimate of $5/2$ by WW14. Year-to-year variations (not shown) do indicate a degree of temporal
 244 non-stationarity.

245 A median spectrum, with best fit of $C = 5.4$, $L_x = 1.2 \times 10^5 \text{ km}$, $\alpha = 1.47$, power law ≈ -2 at
 246 high wavenumbers, is shown in Fig. 8. Also shown are the results for a random set of positions.
 247 While the amplitude changes considerably, the shape of the spectrum is quite stable. The very
 248 large value of L_x is consistent with a long zonal structure in the meridional fluctuations.

249 The decorrelation distance of the meridional wavenumber spectrum is computed from the cosine
 250 transform of Eq. (9). An analytic expression for the autocovariance is,

$$R_l(\xi) = \frac{B \xi^{\alpha-2} \pi^{\alpha-\frac{1}{2}} \left[\pi \xi \left(I_{\alpha+1} \left(\frac{\xi}{L_x} \right) - I_{1-\alpha} \left(\frac{\xi}{L_x} \right) \right) + L_x \alpha I_{\alpha} \left(\frac{\xi}{L_x} \right) \right]}{(2L_x)^{\alpha} \Gamma \left(\alpha - \frac{1}{2} \right) \sin(\pi \alpha)} \quad (10)$$

251 where I_q is the modified Bessel function. This expression has a straightforward behavior only
 252 when $\alpha = m/2$, where m is an integer, and involves subtracting growing exponentials. In practice,
 253 a numerical calculation of the cosine transform proves more robust. Generally speaking, the auto-
 254 covariance of both the median, and from the pointwise calculation at most places, produces a zero
 255 correlation at about 750 km. Beyond that distance, the autocovariance often becomes strongly
 256 negative.

257 *Zonal Wavenumber Spectra*

258 Zonal wavenumbers reflect the meridional quasi-geostrophic variability structure. Zonal flow
 259 dominance at long periods suggests, a priori, that the zonal wavenumbers may be much noisier
 260 (unstable) than the meridional ones. Results for the zonal structure parameters are shown in Figs.
 261 9-11 and the instability of results is manifested in the complex spatial variations in C in Fig.
 262 9. Analyzed zonal arc length varies with latitude from about 2000 to 3300 km. As with the
 263 meridional wavenumber values, the zonal wavenumber periodograms were computed for each
 264 location and each year and then averaged over the 20 years, thus suppressing the year-to-year
 265 variability. Parameter values (from fitting the median spectrum) are: $B = 63$, $L_y = 7910$ km,
 266 $\alpha = 1.6 \approx 3/2$, power law, ≈ -2 , again.

267 With the different results for zonal and meridional spectra, any wavenumber isotropy assumption
 268 must be examined carefully and is surely different for time scales associated with internal waves
 269 and balanced and sub-mesoscale eddies than it is for the general circulation scales dominating the
 270 results here.

271 5. Vertical Structure

272 With some exceptions (e.g., Blumenthal and Briscoe 1995; Polzin and Lvov 2011) the inter-
273 nal wave band can be represented with vertically propagating free waves—that is without vertical
274 modal structure. In contrast (Wunsch 1997, Arbic et al. 2014, Hochet et al. 2015, Lacasce 2017),
275 the balanced eddy field band appears to be dominated by the barotropic and first baroclinic modes
276 with higher modes being much weaker. The often-observed coupling of these two lowest modes
277 (Wunsch 1997, Wang et al. 2013, Lacasce 2017, and others) is rationalized as the tendency of
278 topographic features to minimize the horizontal velocity at and near the bottom over topography.
279 These inferences rely heavily on the in situ current meter measurements that have been accumu-
280 lated over the last several decades, and with some indirect inferences made from the altimeter data.
281 The longest in situ records do not generally surpass two years in length, and their geographical
282 distribution remains sparse and irregular. ²

283 Absence of long duration moored data means that little or nothing is known about the full vertical
284 structure of temperature variability over years and decades apart from some regional inferences
285 (e.g., Bindoff and McDougall, 1994) based upon the temporally sparse deep CTD data. In an
286 initial reconnaissance, the ECCOv4 state estimate/climatology will be used to describe the vertical
287 structures extending out to 20 years. How is that best done?

288 In the context of the balanced eddy band the moored data, where available, produced a reason-
289 ably efficient representation in terms of the linear, *flat bottom*, modes (which are a complete set).
290 Such a representation has the virtue of being consistent with the horizontal Fourier frequency/-
291 wavenumber representations of WW14. On the other hand, the inference of very strong topo-
292 graphic effects, and the known complexities in the time and space scales (e.g., Wunsch 2015;
293 Lacasce 2017) of vertical and horizontal propagation of linearized β –dominated motions make

²The isolated Bermuda Station S and the Hawaiian HOT series are exceptions albeit subject to temporal aliasing problems.

294 any such choice seem arbitrary. Consider the differences between linear baroclinic wave propaga-
 295 tion in the equatorial regions, at mid-latitudes, and at high-latitudes. Response e.g., in the Rossby
 296 wave regime, to eastward going forcing disturbances produces vertically trapped (“negative equiv-
 297 alent depth”) solutions, while westward moving forcing radiates wave-like motions into the deep
 298 interior. Varying topographic slopes will have very different effects on energy that does reach
 299 the bottom topography, and the influences of mean flows are very important, particularly at high
 300 latitudes.

301 In the spirit of exploration, we instead here use the singular value decomposition (empirical or-
 302 thogonal functions, or EOFs, and several other terminologies). Consider the temperature anomaly
 303 field, written as a matrix for $\theta(\mathbf{r}_j, z_m, t_n)$ at horizontal position \mathbf{r}_j , at vertical positions z_m , and
 304 times (years t_n ; only yearly averages are being used for this purpose). With \mathbf{r}_j fixed, the Eckart-
 305 Young-Mirsky Theorem shows that a *perfect* representation by the singular value decomposition
 306 (SVD) is,

$$\theta(\mathbf{r}_j, z_m, t_n) = \mathbf{U}\mathbf{\Lambda}\mathbf{V}^T = \sum_{j=1}^L \lambda_j(\mathbf{r}_j) \mathbf{u}_j(\mathbf{r}_j, z_m) \mathbf{v}_j^T(\mathbf{r}_j, t_q), \text{ fixed } \mathbf{r}_j \quad (11)$$

307 where the orthonormal columns, $\mathbf{u}_j(z_m)$, of \mathbf{U} carry the vertical structure and are often known as
 308 the EOFs. Arguments z_m and t_q have been written on the right side of Eq. (11), for mnemonic
 309 reasons. Each \mathbf{u}_j has length equal to the number of local depths being used, and the number of
 310 columns, L , in \mathbf{U} , is less than or equal to the smaller of the number of depths, z_m and the number
 311 of years, here 20. Thus L never exceeds 20. The orthonormal columns, \mathbf{v}_m , of matrix \mathbf{V} each
 312 carries the time representation of the corresponding \mathbf{u}_m . Their length is the number of years.
 313 Diagonal matrix $\mathbf{\Lambda}$ has elements λ_j known as the singular values. $\mathbf{u}_j, \mathbf{v}_j$ are the singular vectors.
 314 (See Lawson and Hansen 1995 or Wunsch 2006; $\mathbf{\Lambda}$ is not usually square, but a main diagonal is
 315 still defined in the obvious way.) Dimensions are most conveniently associated with λ_j , the vectors
 316 being dimensionless.

317 This representation is particularly useful when the *effective* number of singular values/vectors is
 318 very small compared to the maximum number possible. If Eq. (11) is truncated at a value $K \leq L$,
 319 the fraction of the variance in θ that is captured by the representation is,

$$F(K) = \frac{\sum_{j=1}^K \lambda_j^2}{\sum_{j=1}^L \lambda_j^2} \leq 1, K \leq L \quad (12)$$

320 If a useful value of K is much less than L , e.g., 1 or 2, a concise description is available of an
 321 otherwise potentially very complex field. A corollary of the Eckart-Young-Mirsky Theorem is
 322 that no other pair of K -orthogonal vectors can increase the captured variance. In what follows, the
 323 vertical and temporal structure of the annual mean temperatures in the state estimate are explored,
 324 with a focus on those regions where a very small number (i.e., $K = 1$ or 2) captures the structure
 325 of interannual variability.

326 A priori, the barotropic mode, if it exists, is not expected to make a measurable contribution
 327 to temperature variability. The variance of temperature is such a strong function of depth that
 328 fitting temperature directly produces results dominated by the upper ocean. For numerical accu-
 329 racy, $\theta(\mathbf{r}_h, z)$ is first weighted by dividing the temperature anomaly by the *local* time average full
 330 vertical temperature profile, $\partial \langle \theta_0 \rangle / \partial z$, as determined from the state estimate:

$$G(\mathbf{r}, z_j, t) = \frac{\theta(\mathbf{r}_j, z_m, t_q)}{\partial \langle \theta_0(\mathbf{r}_j, z_m) \rangle / \partial z} \quad (13)$$

331 The SVD is thus being applied to a weighted temperature, rendering the units of the singular
 332 values as meters. Because $\partial \langle \theta_0 \rangle / \partial z$ is itself noisy, the effect of the division in Eq. (13) is to make
 333 the displacement even noisier. Experiments were conducted by using instead the global median
 334 value of $\partial \langle \theta_0 \rangle / \partial z$, but although it produced a somewhat smoother result, the variability in the
 335 mean profile is sufficiently great spatially that use of a fixed form can distort the results. G should

336 *not* be interpreted as a vertical displacement, as that requires the validity of the local equation,

$$\frac{\partial \theta}{\partial t} + w \frac{\partial \theta}{\partial z} = 0 \quad (14)$$

337 which has no contribution from diffusion or from lateral advection, $\mathbf{u}_h \nabla_h \theta$. Both processes are
338 important at periods of years and longer. G is best interpreted simply as a vertically weighted
339 temperature anomaly.³ The vertical structure of temperature itself is recovered by multiplying back
340 $\mathbf{u}_j(\mathbf{r}_j, z_m) \partial \langle \theta_0(\mathbf{r}_j, z_m) \rangle / \partial z$, re-weighting to the upper ocean. The only drawback compared
341 to applying the SVD to unweighted temperatures is that these re-weighted \mathbf{u}_j are not mutually
342 orthonormal. Fits here are made between 105m depth, and 3900m depth or the bottom, whichever
343 is shallower. Unlike the balanced eddy band, the only qualitative generalization is that the result
344 is spatially complex. Examples along two longitudes are shown in Fig. 12. Regions where the
345 lowest singular vector, $\mathbf{u}_1(z)$, contain 90% or more of the variance (Fig. 13) appear to be restricted
346 largely to the South Atlantic and Indian Oceans. Use of the first two \mathbf{u}_i results in a much greater
347 coverage (Fig. 14), although the corresponding vertical structures in \mathbf{v}_i are very diverse. Any
348 relationship to the linear flat bottom modes, or e.g., the “surface modes” of Lacasce (2017) and
349 others, and their relevance at these much longer periods, remains to be explored.

350 The conventional flat bottom baroclinic vertical displacement mode number is the number of
351 zero-crossings plus one (Fig. 15). Thus the first baroclinic mode represents a unidirectional move-
352 ment of the whole water column up or down and the 2nd vertical mode has one zero crossing at

³A global test (not shown) of the time-average “abyssal recipes” point balance of $w \partial < \theta_0 > / \partial z = k \partial^2 < \theta_0 > / \partial z^2$ showed only extremely limited regions of useful accuracy, even in regions of relatively flat time-mean isopycnals. Terms such as $u \partial < \theta_0 > / \partial x$ are important and that is likely also true of the time-dependent balances. Liang et al. 2017) discuss the complexity of the calculated w field, and the importance of the bolus contribution.

353 depth.⁴ A unidirectional mode (no zero crossings) does dominate much of the Pacific and Indian
354 Oceans, but with considerable regions having a sign reversal with depth. One might have hoped
355 that the \mathbf{u}_i would reflect local linear dynamical modes, but given the general inaccuracy of Eq.
356 (14), any such interpretation should be resisted. Instead, a complicated vertical structure emerges
357 even where one singular vector is dominant, particularly in the South Atlantic and northern Indian
358 Oceans.

359 Although not shown here, the same calculation using the unweighted temperature anomalies has
360 a similar vertical complexity, although the magnitudes in the abyssal ocean are much smaller, as
361 expected, and with deep reversals with depth.

362 **6. Temporal Trends**

363 Each vertical orthonormal singular vector \mathbf{u}_i is accompanied by a time-varying orthonormal
364 vector \mathbf{v}_i . The temporal variations of the lowest, singular vector, \mathbf{u}_1 often exhibit a visual trend
365 over the duration of the state estimate. The significance of these trends, in the presence of a
366 general rednoise in the variability, the known long memory in the ocean, and concerns about
367 residual model drifts make it not easy to evaluate. Model drift is however, suppressed by the use
368 of data sets spanning the whole time domain.

369 Ultimately, a breakdown of the \mathbf{v}_i by frequency band is sought. But given the brief 20-year
370 interval available, only the time-domain structures are described. By fitting a least-squares straight
371 line to each \mathbf{v}_1 , about 80% of the results are significant at two-standard deviations (determined
372 from the fit itself; Fig. 16) and a near-Gaussian distribution. A number (far from all) of the
373 \mathbf{v}_1 visually depict a quasi-linear trend (e.g., Fig. 12). Of those, despite the largest value at the

⁴This description is incomplete in the sense that very close to the sea surface, another reversal of sign occurs, one providing movement of the free surface to compensate the upper ocean thermally derived pressure field. (See Wunsch 2013). By starting the fits below 100 meters, this structure is unseen by the SVD analysis.

374 maximum positive slope, slightly more than half the significant slopes are negative, but the net
 375 heating or cooling can be calculated only from the full water column (see references in Fukumori
 376 et al. 2018). Regions of positive and negative slope are shown in Fig. 16.

377 Some of the intricate structure of trends and depth dependence can be seen in Fig. 12. In the At-
 378 lantic, the temporal change at all but the highest latitude shown (42°N) is a nearly monotonic trend
 379 (warming). That warming corresponds to those depths where the corresponding \mathbf{u}_1 is positive, but
 380 would be a cooling where \mathbf{u}_1 is negative.

381 In the east central Pacific Ocean, a cooling trend is visible at 21°N, predominantly at the
 382 bottom—consistent with existing analyses showing deep North Pacific cooling (Fukumori et al.
 383 2018; Gebbie and Huybers 2019). At 30°S \mathbf{v}_1 is nearly unchanged until about 2008 when a warm-
 384 ing over almost the whole water column sets in. Near the equator (0.6°N), a maximum at the time
 385 of the 1997-1998 El Niño is visible and is a general feature of the near equatorial Pacific Ocean.

386 7. Other Condensed Descriptors

387 The central difficulty encountered in this description is the apparent need for a localized
 388 frequency/wavenumber/covariance decomposition at every point with the volume of numbers be-
 389 coming indigestible. One possible condensation is given by the decorrelation times or distances
 390 discussed above. Other summary numbers, useful for model comparisons, etc. do exist: for ex-
 391 ample, the spectral moments. Consider e.g., the frequency moments. Vanmarcke (1983) defines
 392 the “characteristic frequency of the spectral moments” μ_k of a power density spectral estimate,
 393 $\Phi(\omega')$,

$$\mu_k = \left[\frac{\int_0^{\omega'_{\max}} \Phi(\omega') \omega'^k d\omega'}{\int_0^{\omega'_{\max}} \Phi(\omega') d\omega'} \right]^{1/k}, \quad \omega'_{\max} = 1/(2\Delta t). \quad (15)$$

394 These have dimensions of a frequency. Of particular importance are μ_1, μ_2 which can be interpreted
 395 as the mean frequency, and the root-mean square frequency. Another useful summary number is

396 the “spectral bandwidth” (Vanmarcke, 1983) $\varepsilon = (1 - \mu_2/\mu_4)^{1/2}$. Fig. 17 shows the calculated
397 value of the period $1/\sqrt{\mu_2}$ from numerically summing the frequency spectrum derived from the
398 *annual mean* values. A useful physical interpretation is (Vanmarcke 1983, Eq. 4.4.6) that the ex-
399 pected rate of zero-crossings of the temperature anomaly at each point is just $\sqrt{\mu_2}$ and thus $1/\sqrt{\mu_2}$
400 is an estimate of the interval between sign changes in the temperature anomaly through time. See
401 Wunsch and Heimbach (2013) for oceanographic application to the occurrence of extreme events
402 and other properties. The calculated range here is roughly by a factor of two, with a general,
403 if not consistent, tendency for low latitudes to have shorter intervals than higher latitudes and a
404 complicated behavior near the western boundary currents.

405 **8. Summary and Discussion**

406 Results here are a tentative description of low frequency variability, subject to the numerous
407 assumptions made throughout, and are a mixed bag. The “why” has to be taken up separately.
408 To a great extent, the best qualitative global description is the characteristic complexity itself. On
409 the other hand, a simplified form of the Wortham and Wunsch 2014; WW14) temporal frequency
410 spectral form does succeed in describing with useful accuracy much of the near-global temperature
411 variability in the ECCOv4 state estimate. Dominant zonality of oceanic low-frequency variability
412 emerges in the spatial stability/instability of the meridional/zonal wavenumber spectra. Vertical
413 structure corresponding to the lumped variability at periods lying between 1 and 20 years proves to
414 be complex, vertically, horizontally, (and temporally), and without any obvious globally dominant
415 physics. The problem appears to be somewhat like that faced by geologists: in that field, every
416 location has a nearly unique, noisy stratigraphy and geochemistry, but regional syntheses exist,
417 and a few principles (those of plate tectonics, volcanism, mountain building, sedimentation, etc.)
418 can be perceived as acting globally with widely varying importance.

419 With the numerous long time-scales of adjustment/memory in the ocean, and the complicated,
420 effectively permanent, topographic features, that a 20-year interval produces a very complex pat-
421 tern of variability is not surprising. Oceanic bottom topography does influence the whole wa-
422 ter column—which becomes clearer as near-surface higher frequency structures are suppressed by
423 averaging—and is a very complicated, spatially non-stationary, anisotropic, two-dimensional per-
424 manent disturbance. Over long time scales, the variety of atmospheric variations in space and
425 time, in wind, precipitation, evaporation, and temperature, also affect the underlying ocean in
426 geographically complicated ways, including oceanic flow redistribution effects.

427 In contrast with much variability in nature generally, the structure of the higher frequencies (the
428 balanced eddies and the internal wave band) is simpler here than in the lower frequencies. What
429 remains imponderable is whether e.g., a 200 year duration would produce a simpler pattern of
430 time mean and variability? Some insight into the 100 year and longer changes in the ocean can be
431 found in Roemmich et al. (2012) and Gebbie and Huybers (2019).

432 The reader is reminded that these results all rest on the accuracy of the least-squares-fitted EC-
433 COv4 version of the MITgcm. In particular, apart from temporally and spatially sparse CTD casts,
434 direct measurements of the lower 50 percent of water column temperature variability over recent
435 years and decades are lacking. Inferences here rest upon the ability of the dynamical equations to
436 combine the diverse altimetric, Argo, meteorological, and other data to infer the full water column
437 physics. Extended-duration full-water-column measurements would be useful as tests of system
438 skill.

439 Another important question is whether better methods exist to depict the spatially and tempo-
440 rally changing character of oceanic variability? The appeal and power of Fourier representations
441 is clear, but the non-stationary character of the fields renders awkward conventional results—
442 particularly those applying to the wavenumber domain. Alternatives do exist. Wunsch and Stam-

443 mer 1995) produced a global description of altimetric variability using spherical harmonics. Adap-
444 tive methods, perhaps based upon empirical mode decompositions (Huang et al. 1998) or Slepian
445 functions (Simons et al. 2006) or cluster analysis or some combination need also to be explored.

446 (All of the fields discussed here are available in Matlab form on request to the author and all
447 ECCO fields are publicly accessible through the Jet Propulsion Laboratory.)

448 *Acknowledgments.* Funded in part from the Jet Propulsion Laboratory, NASA Contract
449 NNN12AA01C with MIT. I had very useful comments from Brian Arbic, Joe Lacasce, Raf Fer-
450 rari, Joern Callies and two helpful anonymous referees. Thanks to all the members of the ECCO
451 Consortium and the originators of the data who made this work possible.

452 APPENDIX

453 This Appendix includes some additional charts showing the results of the analytic fits to the
454 estimated frequency and wavenumber fields. For the frequency spectra, Figs. A1, A2 show the
455 amplitude and exponential scale factor of the non-dispersive line (NDL). Amplitudes are generally
456 considerably weaker than in the continuum term. Also shown (Figs. A3, A4) are the exponents α
457 for the meridional and zonal wavenumber spectra and which prove relatively featureless.

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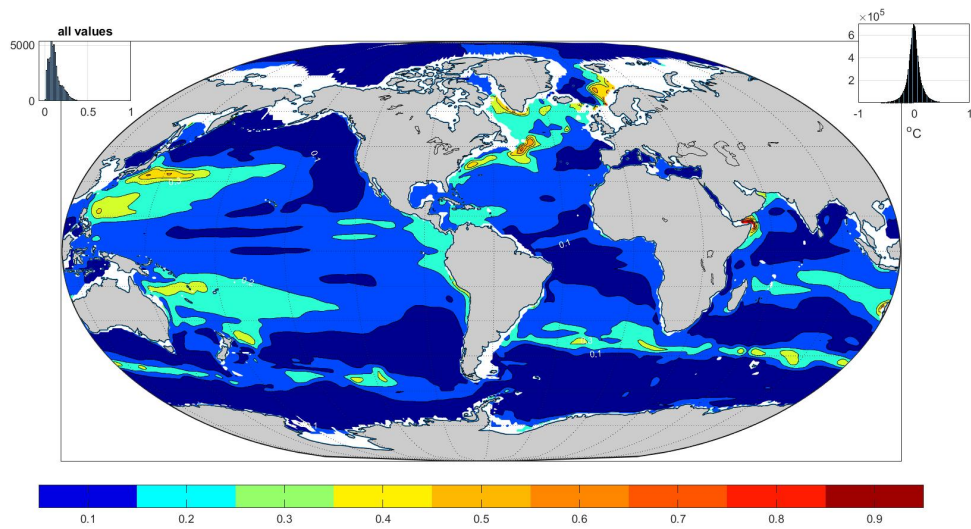
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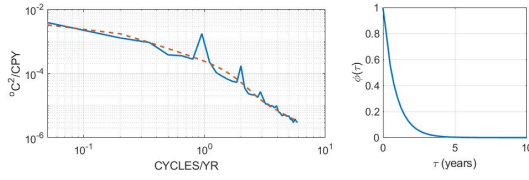
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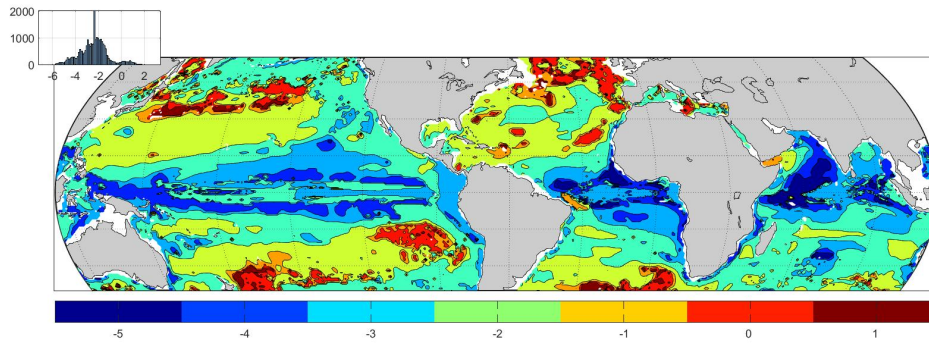
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623	Fig. A4. Estimated value of α from the fit to the zonal wavenumber spectrum. These values are more nearly constant near the median of $\alpha = 1.6 \approx 3/2$.	50
624		



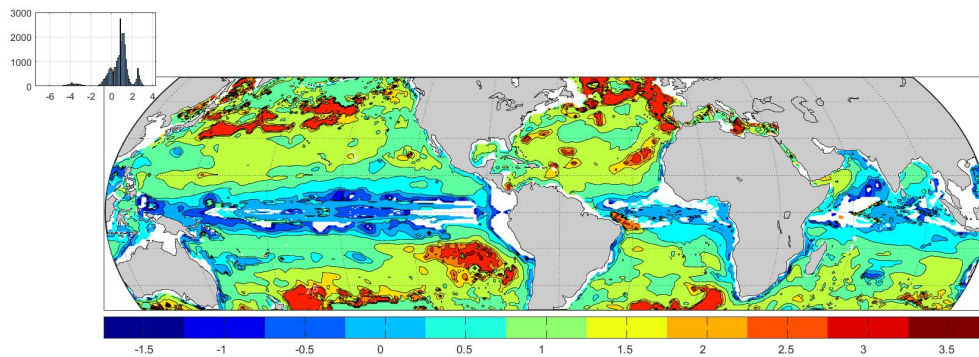
625 FIG. 1. Standard deviation ($^{\circ}\text{C}$) from monthly anomalies of the temperature at 477m. Upper left inset is the
 626 distribution of standard deviations, and upper right shows the distribution of underlying temperature anomalies
 627 over all points and times and its unimodal character. The mean value is 0.12 and the median 0.10 $^{\circ}\text{C}$.



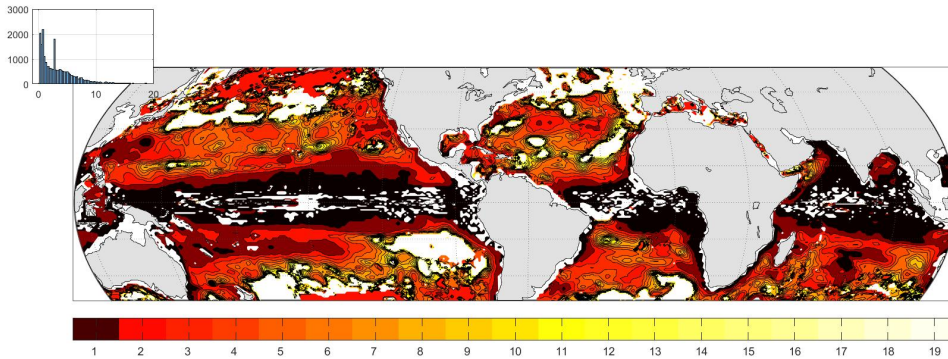
628 FIG. 2. Median spectral density estimate, $^{\circ}C^2/cpy$, from all those computed of the monthly mean temper-
 629 atures at 477m in the state estimate (solid line). Frequency, ω' is cycles/year. Dashed line is the fit of the
 630 analytic spectrum to the median. Note the prominent annual, semi-annual and higher harmonic peaks which
 631 are included in the overall fits. Direct fits to the median spectrum produce $A_1 = 0.058^{\circ}C^2/cpy, T = 8.1y,$
 632 $A_2 = 0.003^{\circ}C^2/cpy, P_a = 1.6y^2$. The decorrelation times and distances discussed in the text would enter into a
 633 formal calculation of the confidence limits on this result. The 95% confidence limit is approximated by about
 634 the small excursion limits in the high frequency range.



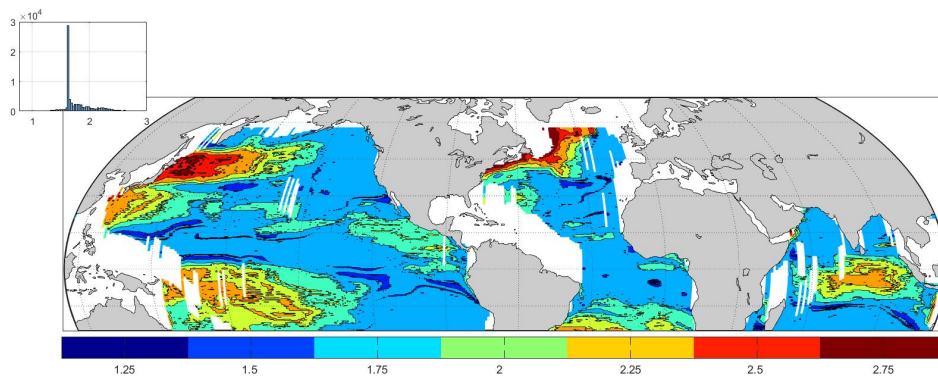
635 FIG. 3. $\log_{10}(A_1)$, where $A_1, ^\circ C^2/cpy$, is the fitted coefficient. A tendency for small values in the tropics and
 636 high values at higher latitudes is evident, albeit with a considerable zonal structure. Compare to Fig. 1 showing
 637 that A_1 is not simply an amplitude, but that the total power is controlled also by T .



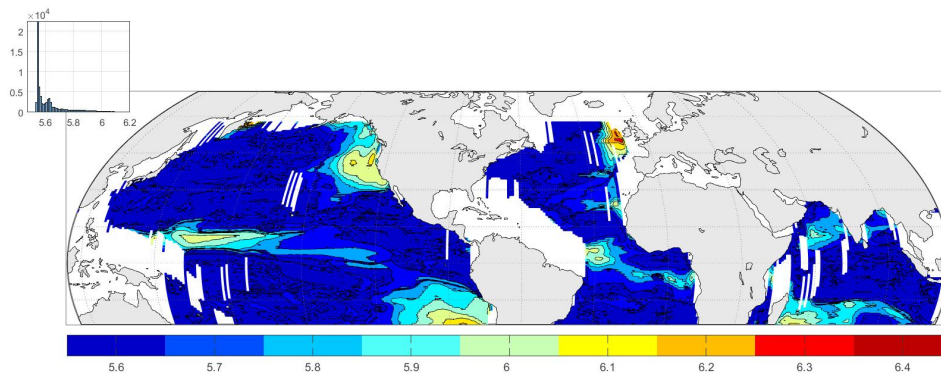
638 FIG. 4. $\log_{10}(T)$ years where T is from the nonlinear least-squares fit. A strong spatial correlation with A_1 is
 639 evident consistent with the behavior of the inverse Hessian (not shown) Note the multi-modal behavior.



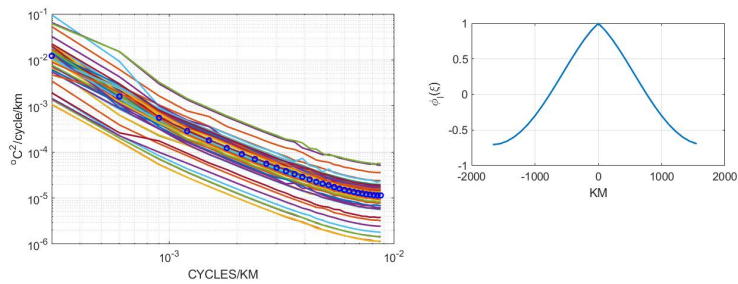
640 FIG. 5. Time in years τ_d for the magnitude of the temporal autocorrelation to fall to 0.1. In white areas, the
 641 correlation never becomes as small as 0.1. Based on the both the continuum ω'^2 and NDL terms in the analytic
 642 form in Eq. (7a).



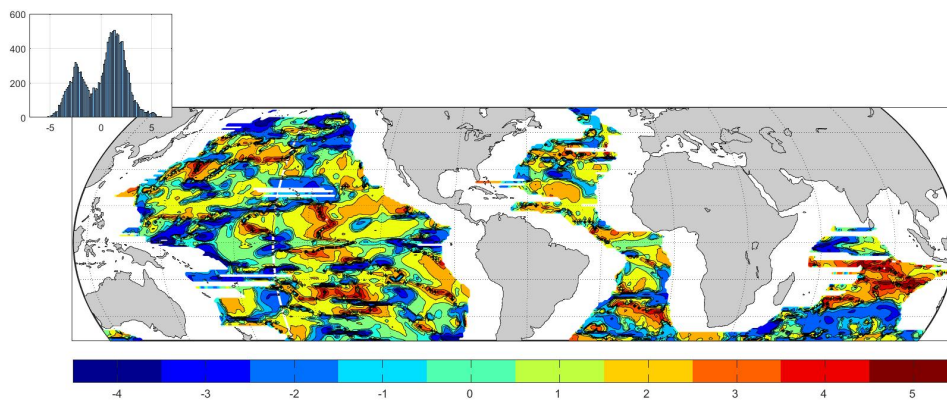
643 FIG. 6. $\log_{10}(C)$, $^o C^2/\text{cpy}$, in the meridional wavenumber spectral density (Eq. 9). Both visually and in the
 644 histogram, a considerable fraction of the ocean has a nearly uniform value, albeit the western boundary areas are
 645 distinctly stronger. Values are assigned to the mid-points of vertical strips of 30 degrees of latitude and white
 646 areas are within 15 degrees of the coast. Isolated white stripes result from island presence in the strip.



647 FIG. 7. The nearly uniform open ocean value of $\log_{10}(L_x)$ in the meridional wavenumber spectrum. Here
 648 eastern boundaries tend to stand out. Median value is about 380km. Mean is about 460km.



649 FIG. 8. (left panel). Median meridional wavenumber spectrum ('o', not the fit) and wavenumber spectra from
 650 a random selection of positions. (right panel) Autocorrelation corresponds to the median meridional wavenum-
 651 ber spectrum and is typical of most of the ocean. The spectrum remains reddish at long wavelengths..



652 FIG. 9. $\log_{10}(B)$, $^{\circ} C^2/cpy$, in the zonal wavenumber spectral density. Note that the 30° choice of zonal
 653 extent precludes values extending westward more than 30° from the coastline. Values are assigned to the mid-
 654 point position in longitude. In a quasi-geostrophic system, zonal wavenumbers reflect the meridional flow field.
 655 The result is bimodal.

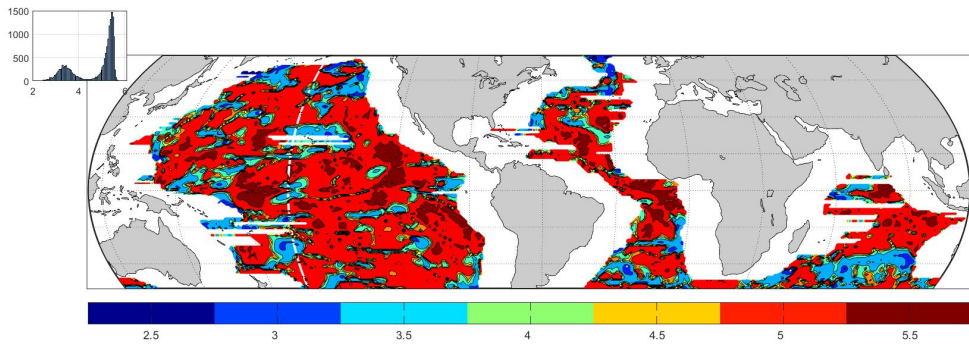


FIG. 10. \log_{10}, L_y in kms, in the zonal wavenumber spectral estimates. Strong bimodality again appears.

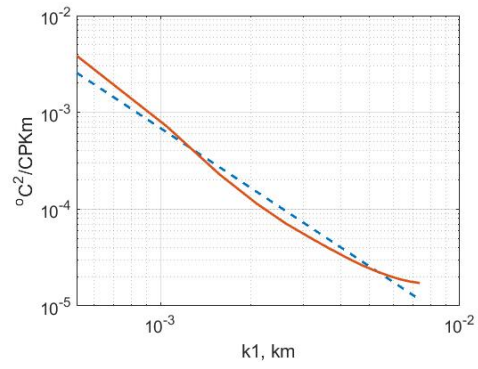
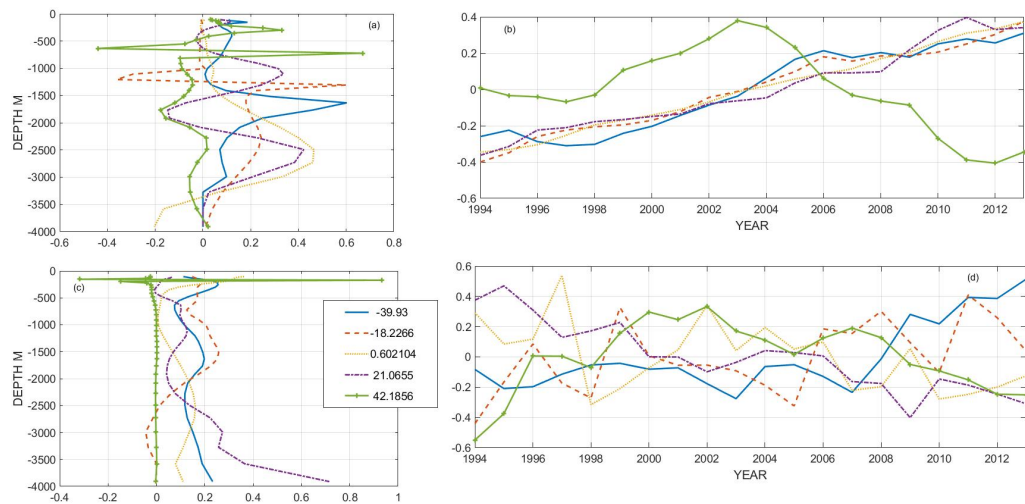
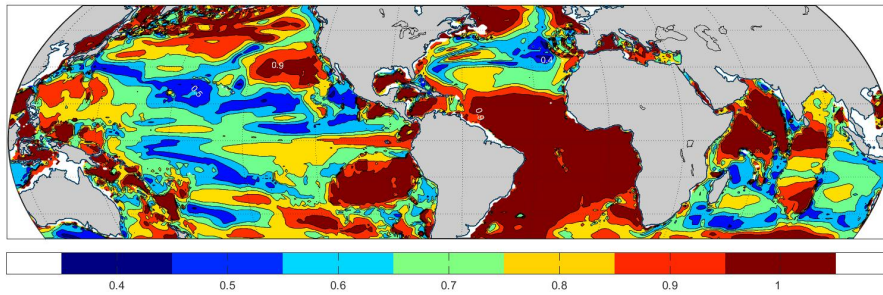


FIG. 11. Median zonal wave number spectral density and the corresponding fit (dashed line).



656 FIG. 12. Pairs of $\mathbf{u}_1(a), \mathbf{v}_1$ (b) along the meridian 10°W in the Atlantic Ocean at every 25th meridional grid
 657 point. Linear trends are visually present, with the strong exception at 42°N , where the initial years imply cooling
 658 (increase of negative values of \mathbf{u}_1) and warming after about 2004. (c,d) Same as (a,b) except along longitude
 659 150°W in the Pacific Ocean. Linear trends are not obvious and note the strong bottom intensification at 21°N of
 660 a slight cooling.



661 FIG. 13. Fraction of the variance lying in the first singular vector (EOF) of vertically weighted temperature
 662 through depth and time (years). Regions where the value exceeds 0.9 are largely limited to the South Atlantic
 663 and the western Indian Ocean. In general the Atlantic is simpler in this special sense than is the Pacific, but the
 664 structures of \mathbf{u}_1 are very variable within those areas.

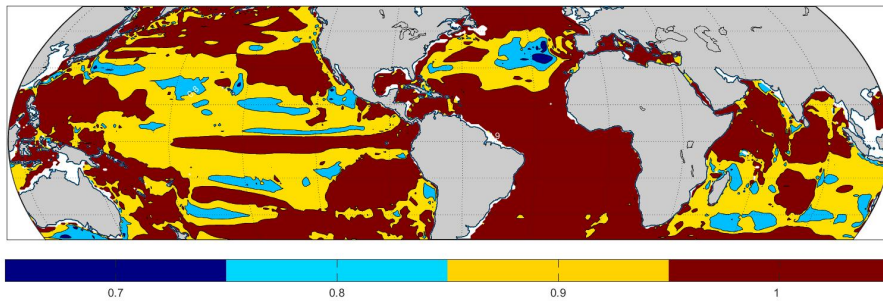
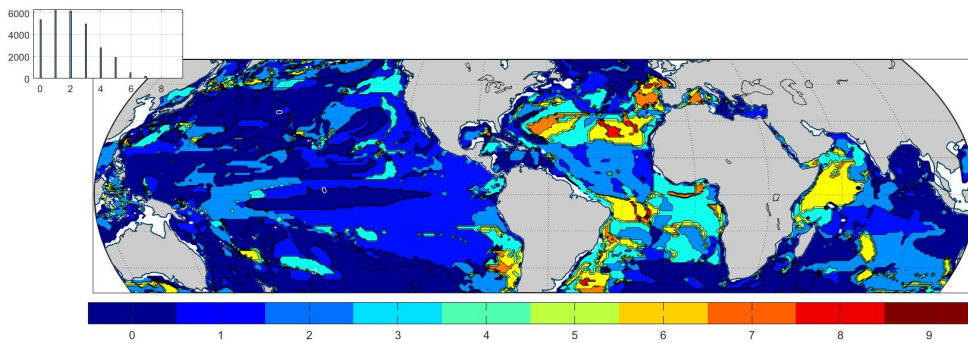
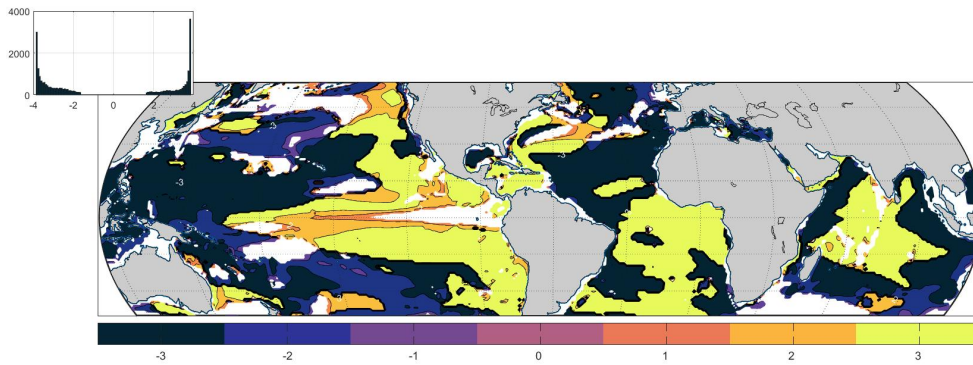


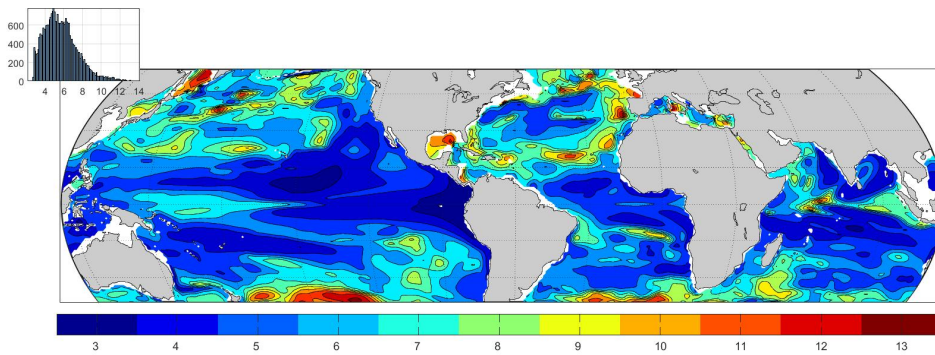
FIG. 14. Fraction of the variance included in the first *two* \mathbf{u}_i singular vectors. Most of the ocean exceeds 90%.



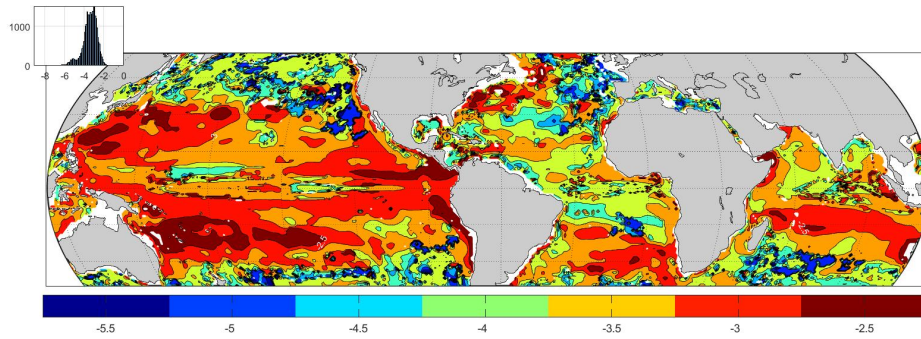
665 FIG. 15. Number of zero crossings, n_z , in the vertical in the first singular vector \mathbf{u}_1) for temperature in
 666 depth/time. In conventional mode terms, the baroclinic mode number is $n_z + 1$. Regions roughly similar to a
 667 simple first baroclinic mode ($n_z = 0$, possibly representing a unidirectional vertical displacement) are unusual.
 668 More common are regions resembling a mode 2 dominance with a zero-crossing in the vertical.



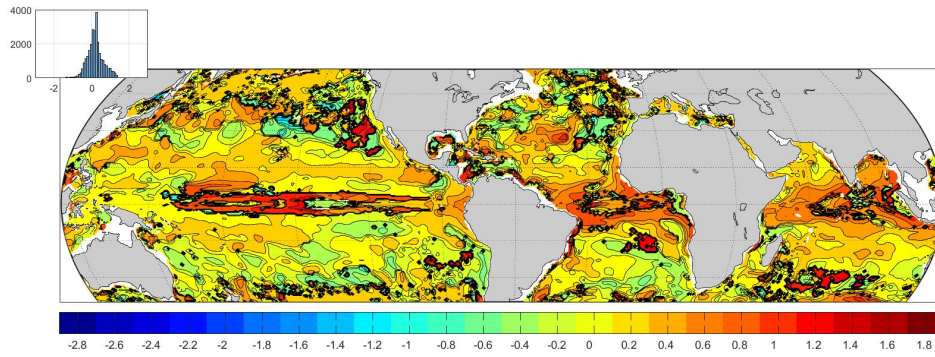
669 FIG. 16. Values of $100\times$ the linear slope in \mathbf{v}_1 where it exceeds 2 standard deviations of the fit error. *The*
 670 *slope pertains to the value of the maximum of \mathbf{u}_1 , defined as always positive. Thus a negative value here means*
 671 *that the depth of maximum change is cooling in the 20 year average. Whether the rest of the water column is*
 672 *cooling or warming depends upon the number of zeros in \mathbf{u}_1 . Some locations show no trend of either sign (white*
 673 *areas).*



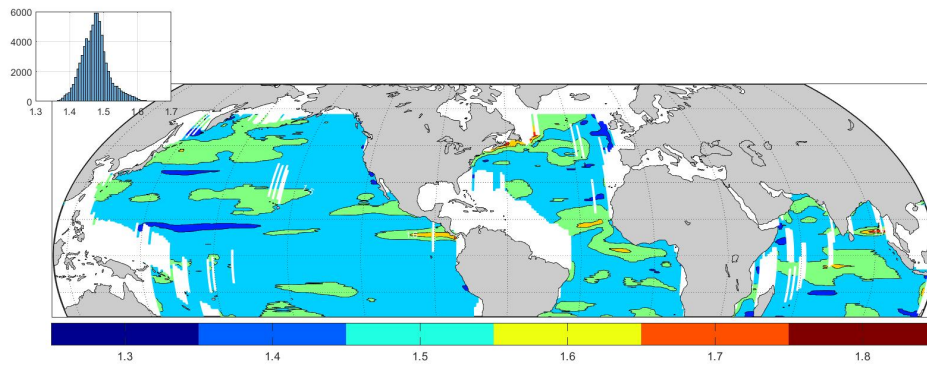
674 FIG. 17. Time scale (years) associated with the second moment of the frequency periodogram at each grid
675 point at 477m annual average values.



676 Fig. A1. $\log(A_2)$, $^{\circ}C/cy$, the coefficient of the nondispersive line frequency contribution. A_2 is generally
677 an order of magnitude smaller than A_1 .

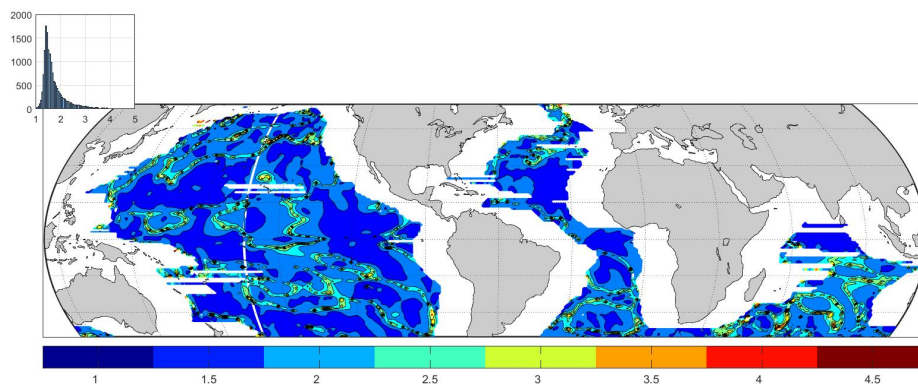


678 Fig. A2. $\log(P_a)$, years², from the fit of the exponential frequency term in the analytic spectrum. A low
 679 latitude and western dominance is the most conspicuous feature along with a high latitude intricacy. Fit is
 680 unstable in the equatorial Pacific and the contouring is incomplete there.



681 Fig. A3. Value α as the best fit in the meridional wavenumber spectrum, and dominated by a value near

682 $\alpha = 3/2$.



683 Fig. A4. Estimated value of α from the fit to the zonal wavenumber spectrum. These values are more nearly
 684 constant near the median of $\alpha = 1.6 \approx 3/2$.