

Ocean Mixing: Oxford Encyclopedia.

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Summary

Oceanic mixing is one of the major determinants of the ocean circulation and its climatological influences. Existing distributions of mixing properties determine the rates of storage and redistribution within the climate system of fundamental scalar tracers including heat, freshwater, oxygen, carbon, and others. Observations have over-turned earlier concepts that mixing rates might be approximately uniform throughout the ocean volume, with profound implications for determining the circulation and its properties. Inferences about past and potential future oceanic circulations and the resulting climate influence require determination of changed energy inputs and the expected consequent adjustment of mixing processes and their influence.

Keywords

Mixing, stirring, balanced eddies, internal waves, oceanic general circulation, sub-mesoscale

Mixing and Stirring

The ocean stores, redistributes, and exchanges with the atmosphere a wide variety of substances essential to the climate system including heat (enthalpy), freshwater, kinetic and potential energies, carbon, oxygen, and nutrients of all sorts. An understanding of climate, and its past and

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19 potential future changes, thus requires a complete depiction of the distribution and redistribu-
20 tion of these scalar properties once they have entered the ocean. In other words, the oceanic
21 general circulation in all of the elements that control it and change it must be described if
22 varying climate states are to be understood.

23 In a laminar flow, physics dictate that the movement of scalar properties is cleanly divisible
24 into advective components—proportional to the velocity field, \mathbf{v} , and diffusive ones dependent
25 upon the molecular diffusion coefficients, κ , arising from the internal energy of the fluid. In
26 a turbulent fluid like the ocean, the distinction is no longer clear-cut: the inability to analyze
27 and cleanly represent the flow field from the largest scales down to the molecular ones has led
28 historically to representations of the unresolved scales as, at least in part, as “diffusion-like.”
29 This situation renders the study of oceanic mixing processes as a subset of a combined laminar-
30 turbulent fluid, but with the turbulence having the special properties of a rotating, stratified
31 fluid shell with complex vertical and horizontal boundaries. A complete discussion of mixing
32 encompasses the entirety of oceanic fluid physics including fully turbulent flows in the presence
33 of stratification, rotation, external forcing, complex boundaries, and global-scale fluid flows.
34 A systematic account of many of these elements can be found in the book by Thorpe (2005).
35 Mixing is simultaneously a consequence of, and a cause of, the large-scale general circulation of
36 the ocean.

37 Several important elements of ocean mixing have been clarified recently: (1) Processes and
38 effective rates vary by three and four orders of magnitude over the ocean volume; (2) The physics
39 of mixing is deeply intertwined with the oceanic energy budget; (3) All of the enormous range
40 of scales of oceanic motions from 10,000 km to 1 mm are both affected by, and in turn influence,
41 variations in mixing. The consequences have proven profound for basic understanding of the
42 oceanic general circulation and its climate impacts. To limit this article, the focus is primarily
43 on the interior ocean, touching only tangentially on the special regions of the mixed-layers at the
44 sea-surface and sea floor, and the variety of boundary layers occurring over sidewalls and general
45 topographic features. Most attention is paid to the vertical or “diapycnal” problem involving
46 transfers across the stable stratification, and which differs fundamentally both in magnitude and
47 physics from so-called isopycnal mixing—easier motions along layers of near-constant density—
48 nearly perpendicular to local gravity.

49 Discussion of the causes and numerical values of mixing, apart from being an extremely
50 challenging observational problem, is at the very edge of modelling and theoretical understanding
51 of fluid physics. Much of the context of this discussion can be understood from the conventional
52 advection-diffusion equation for a tracer, C , written for an ordinary fluid e.g.,

$$\frac{\partial C}{\partial t} + \mathbf{v} \cdot \nabla C = \kappa_C \nabla^2 C. \quad (1) \quad \{\text{advectiondiff}$$

53 κ_C is a Newtonian diffusion coefficient of molecular origin, assumed constant in both space and
 54 time and ∇ is the three-dimensional gradient operator in any suitable coordinate system. The
 55 most common tracers in oceanic studies are temperature, T , and salinity, S . These two fields
 56 determine the density, ρ , structure of the ocean as

$$\rho(\mathbf{r}, t) = \varrho(T(\mathbf{r}, t), S(\mathbf{r}, t), p(\mathbf{r}, t)) \quad (2) \quad \{\text{density1}\}$$

57 where \mathbf{r} is a three-dimensional position ((x, y, z) in Cartesian coordinates) and p is the pressure.
 58 (Variations in ρ are here considered to be “small” in the mass balance, but important in the
 59 momentum equations.) The function in Eq. (2) involves a complicated empirical rule rendering
 60 the density field a nonlinear function of the variables T, S, p . For many purposes, however, it
 61 can be linearized in the form,

$$\rho(\mathbf{r}, t) = \rho_0 (1 - \alpha T(\mathbf{r}, t) + \beta S(\mathbf{r}, t) + \gamma p(\mathbf{r}, t)), \quad (3) \quad \{\text{density2}\}$$

62 where $\rho_0, \alpha, \beta, \gamma$ are treated as locally constant. Often γ is set to 0, in the Boussinesq approx-
 63 imation, treating the fluid as incompressible. If Eq. (3) is adequate, ρ itself can be treated as
 64 a tracer, $C = \rho$.

65 In the simplest cases, \mathbf{v}, κ_C are independent of tracer concentration C (called a “passive”
 66 tracer), but if T, S , or ρ are being considered, they influence \mathbf{v} (“active” tracers). Typical
 67 numerical values are $\kappa_C \approx 10^{-7} \text{m}^2/\text{s}$ for temperature and $10^{-9} \text{m}^2/\text{s}$ for salt. Discussion of
 68 tracer mixing in the ocean is intimately tied to momentum mixing (frictional processes) and
 69 they are best treated together. Thorpe (2007) has a clear introduction. For present purposes,
 70 \mathbf{v} is taken as a “given.”

71 The derivatives in the tracer “diffusion” term on the right-hand side only become important
 72 when the length scales over which they are taken become very small, of order centimeters or less.
 73 An attractive step for an oceanographer attempting to understand the behavior of a tracer in
 74 the ocean is to argue that this diffusion term can be neglected, setting the left-hand-side equal
 75 to zero so that,

$$\frac{\partial C}{\partial t} + \mathbf{v} \cdot \nabla C = \frac{\partial C}{\partial t} + \nabla \cdot (\mathbf{v}C) = 0, \quad \nabla \cdot \mathbf{v} = 0 \quad (4) \quad \{\text{advectiondiff}\}$$

76 As with much of fluid dynamics, most interest and difficulty arises from the overall behavior
 77 of the advection-term, $\mathbf{v} \cdot \nabla C$. Estimates of the spatial variation of velocity, \mathbf{v} , show it carrying
 78 laminar-like structures on the largest scale (e.g., the overall gyre structure of the circulation of
 79 the Pacific Ocean) through an entire stochastic continuum of spatial scales down to 1 millimeter.
 80 As far as can be determined, no spatial scales are absent in the range between these rough limits
 81 (or in the jargon, there is no spatial “spectral gap”). Much, if not all, of the flow contains
 82 powerful random (stochastic) elements—that is, it is turbulent—and is best perceived from

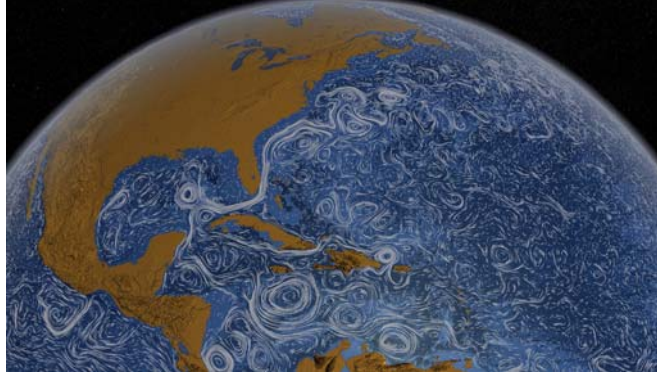


Figure 1: Surface speeds in part of the North Atlantic Ocean (from website <http://www.nasa.gov/topics/earth/features/perpetual-ocean.html#.Vy0DIoSDFBd>). from a numerical model. Most of these features are intensely time-dependent and can be seen in an animation at the same web location. Myriad yet-smaller scale features exist in the ocean that are invisible to the eye in such an overview and remain beyond existing computing capacities.

{perpetual_oce

83 one of the many animations of the ocean circulation now available on the web. Fig. 1 shows a
 84 snapshot of estimated surface speed in the ocean. The visually conspicuous “swirling” flows carry
 85 tracer properties with them, stretching their boundaries out to enormous lengths. These motions
 86 are commonly referred to, somewhat vaguely, as the oceanic “mesoscale eddies”, or “balanced
 87 eddies”. The latter terminology emphasizes their tendency to satisfy near-geostrophic balance
 88 (dominated by the Earth’s rotation) although at the shortest scales, below about 10km, they
 89 become significantly non-geostrophic and are then commonly labelled “sub-mesoscale” eddies.
 90 Below visual detection, molecular processes are also at work, preventing the structures from
 91 becoming arbitrarily elongated.

92 Suppose a patch of dye with a circumferential distance of order D_0 is introduced into such a
 93 flow in two-dimensions (Fig. 2). An the initial time, at the patch edge, ∇C , and hence $\kappa_C \nabla^2 C$,
 94 may not be negligible, but the initial $D = D_0$ is sufficiently small that an integral of the last
 95 term in Eq. (1)

$$\oint_{D=D_0} |\kappa_C \nabla^2 C| dD, \tag{5} \text{{integral1}}$$

96 taken around its boundary, *is* negligible compared to other integrated term from Eq. (1)

97 Through time, the patch is stretched out, and extended in a tortuous pathway. Even in
 98 comparatively simple, non-turbulent flows, pathways commonly have a chaotic behavior (see
 99 e.g. Ottino, 1989). The boundary demarcating the edge of the path becomes extended in time,
 100 growing to values $D \gg D_0$. At this boundary, the gradients, ∇C , are extremely large, and
 101 the integrals in Eq. (5) taken around the long boundary D are very important and cannot be

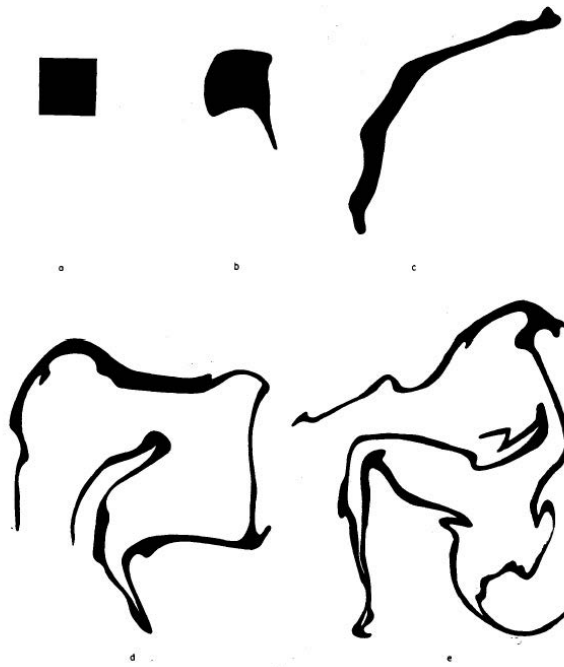


Figure 2: From Welander (1955) showing the “stirring” of a small square dye patch in a near two-dimensional fluid flow. “Mixing” would take place along the dye boundary which becomes greatly extended in length as time goes on. Molecular processes operate primarily at the dye-patch boundaries and ultimately limit the extensions of the patches. In three-dimensions the tendency to elongation is muted.

{welander_1955

102 ignored and indeed prevent the fluid from developing arbitrarily elongated and tangled struc-
 103 tures. Indefinite growth is prevented both by these molecular processes, and those acting to
 104 control particle separations within a dye patch (Salmon, 1998, Section 1.14). Thus one must
 105 distinguish the statistics of turbulent motion insofar as they affect the center of mass of the dye
 106 patch, from those controlling the separation of fluid particles within the patch (see especially,
 107 Garrett, 2006).

108 Eckart (1948) divided oceanic processes into three parts—supposing there exists a large-
 109 scale flow carrying tracer patterns of interest. Superimposed upon those large-scale, L , flows are
 110 smaller scale motions, L_s , which stretch the tracer out into the complex patterns with intricate
 111 and extended boundaries. He called those motions “stirring.” Finally, the very smallest scales
 112 at both patch edges, and within the interior of the patch, the molecular scale l_m , rarely observed
 113 directly, are where molecular processes dominate and act to homogenize the fluid. (Note that
 114 Fig. 2 depicts an essentially two-dimensional fluid movement; in three dimensions the elongated,
 115 serpentine, structure is less-pronounced.)

116 Although the molecular scale was early on recognized as essential to understanding the

117 distribution of properties in the oceans, the observational tools available before about 1965
 118 did not permit discussion of how these mixing processes actually operated. A great deal had
 119 been learned about the physics of conventional turbulence in unstratified, non-rotating fluids,
 120 but the oceanic applicability remained obscure. An oceanographic exception was the growth of
 121 interest in so-called double-diffusive processes, although most of that discussion was focussed on
 122 laboratory-scale experiments and theory (e.g., Radko, 2013).

123 For roughly the first 100 years of physical oceanography (to about 1965), observational
 124 capabilities on the largest scales were limited to ship-borne measurements of temperature and
 125 salinity, resulting in the now-familiar and classical charts of oceanic gyres and major boundary
 126 currents, and are known to be quasi-permanent features. These are identified with scale L , see
 127 Fig. 3, and which were and are represented as a large-scale, fundamentally laminar flow. On
 128 the other hand, oceanographers were well aware of the existence of smaller scales of flow (see
 129 Helland-Hansen and Nansen, 1909) and realized that they had both a dynamical and kinematical
 130 influences on the larger scales, ones that ultimately had to be treated statistically.

131 Assuming that these motions were turbulent, in the sense of being stochastic, and in a
 132 fluid-dynamics approach dating back to the 19th Century, it was hypothesized that both \mathbf{v}, C
 133 could be broken up into two pieces: $\mathbf{v} = \bar{\mathbf{v}} + \mathbf{v}'$, $C = \bar{C} + C'$, where the over-bar represents an
 134 averaging process, in space, or time, or both. If space was involved (a useful choice in a spatially
 135 homogeneous flow) the averaging interval was taken to be larger than the stirring scale L_s , but
 136 less than L , such that $\overline{\mathbf{v}'} = 0$, $\overline{C'} = 0$, but $\overline{\mathbf{v}'C'} \neq 0$. (See especially Tennekes and Lumley (1972)
 137 or any book on turbulence, e.g. Batchelor, 1953). Alternatively, the assumption might be made
 138 that a time average would remove all of the small scales—although that involves a strong and
 139 difficult-to-justify assumption about space-time statistics of the ocean. In particular, the longest
 140 available records do not support the existence of any kind of spectral gap in frequency.¹ Eq.
 141 (4) then becomes, when averaged,

$$\frac{\partial \bar{C}}{\partial t} + \bar{\mathbf{v}} \cdot \nabla \bar{C} + \nabla \cdot (\overline{\mathbf{v}'C'}) = 0, \quad \nabla \cdot \mathbf{v}' = 0. \quad (6)$$

142 $\bar{\mathbf{v}}, \bar{C}$ are the quasi-laminar fields with everything else to be handled statistically. The problem—
 143 the “turbulence closure”—is what to do with the divergence of $\overline{\mathbf{v}'C'}$, the covariance of the

¹Davis (1994) emphasized, absent a spatial or temporal gap in scales, the assumption that such mixed averages such as $\overline{\mathbf{v}'C'} = 0$ can fail. Few oceanic measurements exist such that averages taken over any long time-scale are truly stable: e.g., doubling the record length usually measurably changes the calculated averaged and which may well be covarying with the variability of the fluctuations. The trough in spectra at frequencies just below the inertial frequency (usually denoted f) is sometimes regarded as representing a spectral gap, but it is probably best regarded as the effect of a peak, rather than a minimum. Recent theoretical arguments have also shown how energy can be transferred across this frequency interval.

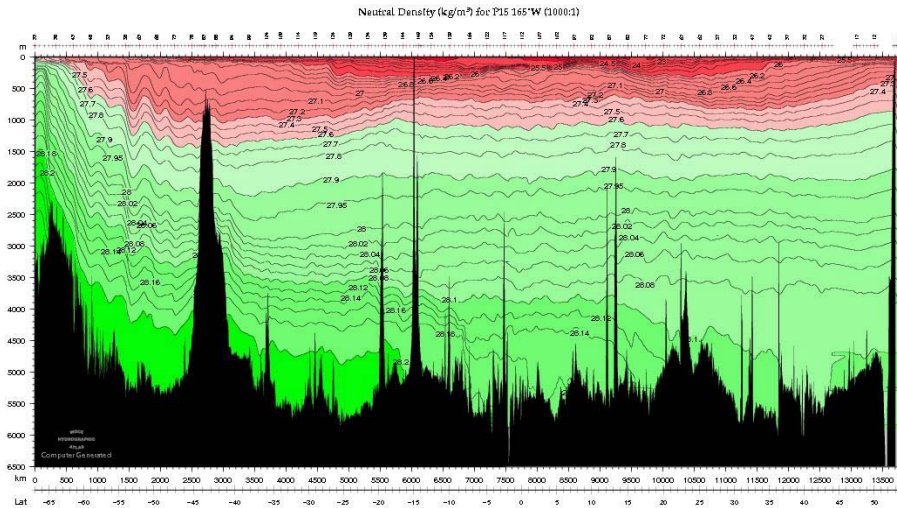


Figure 3: Neutral surfaces down the central Pacific Ocean (from the Pacific Atlas of the World Ocean Circulation Experiment, Talley et al., 2007). Note that the contouring interval is not uniform. The largest scales are quasi-stable over decades, whereas the visual much smaller scale variability was poorly understood until recently, and is best regarded as a mixture of geostrophically balanced motions and gravitationally governed internal waves. Surfaces above roughly 1000m in the Southern Ocean outcrop to the surface before they reach the Antarctic continent. These outcropping surfaces can be stirred and mixed by direct wind-action. At greater depths, surfaces away from the Southern Ocean are roughly horizontal sometimes leading to depictions as the result of one-dimensional, vertical, advection-diffusion balances. These surfaces reach the continent to the south where they become involved with the formation of very dense water (Antarctic Bottom Water) visible as the deep green plume at the bottom left. This water must ultimately be returned to the sea surface if the ocean is in a steady-state, through processes requiring work to mix the properties upward against gravity. The “thermocline” is the region near the surface of the most rapid density change (controlled largely by temperature). Such figures have a very great vertical exaggeration with the horizontal range being nearly 10,000 km and the vertical range being 6 km and thus the topographic and isopycnal slopes are in practice very small, but nonetheless extremely important in many locations. The structure of isopycnal surfaces is very similar to that of neutral surfaces.

{woce_atlas_p1

144 spreading scales? (In conventional homogeneous, isotropic turbulence, this divergence vanishes.)
 145 Following Osborne Reynolds and many others, it can be supposed that these spreading scales
 146 act as an analogue of the suppressed diffusive ones—but operating entirely locally on the L -scale
 147 so that,

$$\nabla \cdot (\overline{\mathbf{v}'C'}) = -K\nabla^2\bar{C}. \quad (7)$$

148 In a turbulent flow, it is often argued that all tracers are stirred by eddies at the same rate
 149 so that K would be appropriate for all C . Because of known anisotropies in oceanic flows and
 150 features, anisotropic rules were used,

$$\nabla \cdot (\overline{\mathbf{v}'C'}) = -K_{xx}\frac{\partial^2\bar{C}}{\partial x^2} - K_{yy}\frac{\partial^2\bar{C}}{\partial y^2} - K_{zz}\frac{\partial^2\bar{C}}{\partial z^2} \quad (8)$$

151 (in a Cartesian system) or in the most general form,

$$\nabla \cdot (\overline{\mathbf{v}'C'}) = -\nabla \cdot (\mathbf{K}\nabla\bar{C}), \quad (9) \quad \{\text{tensor1}\}$$

152 where

$$\mathbf{K} = \left\{ \begin{array}{ccc} K_{xx} & K_{xy} & K_{xz} \\ K_{yx} & K_{yy} & K_{yz} \\ K_{zx} & K_{zy} & K_{zz} \end{array} \right\} = \left\{ \begin{array}{ccc} K_{11} & K_{12} & K_{13} \\ K_{21} & K_{22} & K_{23} \\ K_{31} & K_{32} & K_{33} \end{array} \right\}, \quad (10) \quad \{\text{tensor2}\}$$

153 has become a tensor (written above both in the special Cartesian and a more general index
 154 forms) with spatial or even temporal structure whose physical form must obey certain rules (e.g.
 155 Griffies, 2007, P.288+).

156 Use of these “eddy” coefficients (and their analogues for viscosity in the momentum equa-
 157 tions) have had a long and useful history. Extensive discussions can be found in Defant (1961,
 158 P. 107+). In both analytic and numerical modelling, the symmetric elements of \mathbf{K} destroy any
 159 gradients long before they reach the l_m scale, and thus function *implicitly* to dissipate tracer
 160 variance. That mixing depends only on the *local* properties of C has been assumed: for a
 161 contrasting approach, see Canuto et al. (2007). Anti-symmetric elements can act as stirring
 162 motions, increasing the gradients; see e.g. Griffies (2004) or Flierl and McGillicuddy (2007).

163 At the time of this writing, a major field experiment, Diapycnal and Isopycnal Mixing Ex-
 164 periment in the Southern Ocean (DIMES; Sheen et al., 2013), is underway to better understand
 165 the structure of \mathbf{K} , albeit the physics of the Southern Ocean flow are often untypical of the
 166 global ocean.

167 **Isopycnals and Neutral Surfaces**

168 Most of the ocean is stably stratified (Fig. 3), and the fluid has difficulty in moving tracers across
 169 that stratification, which requires work against gravity. In contrast, movement perpendicular to

170 gravity involves much less work and energy (except for internal viscous stresses), and substantial
 171 evidence supports the inference that usually $K_{33} \ll K_{11}, K_{22}$. In practice, the surfaces of con-
 172 stant density, $x_3 = \sigma(x_1, x_2)$ in many places deviate from the horizontal (Fig. 3 at the southern
 173 end), and the expression representing vertical diffusion can become inaccurate. K_{33} is better
 174 expressed as $K_{\sigma\sigma}$ (a “diapycnal” diffusion) and K_{11}, K_{22} and the various cross-terms become ex-
 175 pressions for diffusion along the isopycnal surfaces (“isopycnal” diffusion). Because the equation
 176 of state of seawater, representing the density as a function of temperature, salinity, pressure, and
 177 location, $\rho(T(x_1, x_2, x_3), S(s, x_2, x_3), p(x_1, x_2, x_3))$ is significantly non-linear, many authors re-
 178 place isopycnals with “neutral surfaces” along which the work done against gravity by movement
 179 along the surfaces is minimal. To emphasize that the “vertical” coordinate will differ from the
 180 geocentric vertical, the coordinate system $(x, y, z) \rightarrow (x_1, x_2, x_3)$ where it is understood here
 181 that x_3 is normal to an isopycnal or neutral surface and is nonetheless sometimes most simply
 182 written as z . Terminology “isopycnal” will be used here without distinguishing the various rep-
 183 resentations of surfaces of minimal work against gravity. K_{33} becomes a “diapycnal” coefficient,
 184 and $K_{11,22}$ are “isopycnal” coefficients. Over much of the ocean, the distinction between the geo-
 185 detic vertical and the isopycnal normal direction can be neglected, but the difference is crucial
 186 in regions of isopycnal slope as seen in Fig. 3 in the Southern Ocean. The non-linearity of the
 187 equation of state gives rise to a number of sometimes puzzling behaviors of mixed fluids under
 188 the rubrics of “cabbelling” and “thermobaricity.” (For example, two fluid parcels with different
 189 temperature and salinity but the same density will, when mixed, have a different density. See
 190 e.g., McDougall and Garrett, 1992; Schanze and Schmitt, 2013.)

191 By definition, in an isopycnal coordinate system, mixing cannot move fluid parcels of different
 192 density across the reference surfaces, but that is not true of passive tracers. This distinguishing
 193 behavior leads to a lot of detailed complexity in numerical models of the ocean.

194 Scales

195 Oceanic flows do not occupy discretely separable spatial scales, L, L_S, l_m , but consist instead of
 196 a continuum of near-geostrophic global horizontal scale structures (basin and larger), changing
 197 significantly in time only slowly (many decades and longer); a geostrophically balanced eddy
 198 field, with spatial scales of a few hundred kilometers and changing over many months and longer;
 199 an internal-inertial wave field on spatial scales of tens of kilometers and changing over hours
 200 to days; a “sub-mesoscale” where rotation no longer dominates the eddies, and a small-scale
 201 turbulence with features of tens of meters to millimeters tending to be three-dimensionally
 202 isotropic, and changing over seconds to days. These motions overlap in a continuous occupation
 203 of frequency/wavenumber spaces. Derivatives of tracers and velocity, taken over the smallest

204 scales usually greatly exceed those taken on the largest scales.

205 A small extension of Eckart’s (1948) schematic description of oceanic mixing is that tracers
 206 varying on scale L are stirred by the balanced (geostrophic) and sub-mesoscale eddies (scale
 207 L_{BE}). Then localized breaking by the internal wave field (L_{IW}) generates the turbulence that
 208 is dissipated in a traditional turbulent cascade to the molecular scale l_m , dominating the fi-
 209 nal conversion into molecular scales. Although a useful conceptual structure, it can be very
 210 inaccurate and misleading. For example, the gyre scales contain many features such as the west-
 211 ern boundary currents, the equatorial current system, etc. which lie on the same scale (order
 212 100km) as do the balanced eddies. Balanced eddies appear to exist even at very long periods—
 213 much longer than a year. They include such diverse physics as linear and non-linear Rossby and
 214 Kelvin waves, isolated vortices, coupled waves, instability structures, etc. Internal waves overlap
 215 the high wavenumber part of the submesoscale eddy field, and also contain balanced “vortical”
 216 structures. This listing remains only a sketch, ignoring many different boundary layer types
 217 at the surface, sidewalls and bottom. All of these physical processes are coupled to a greater
 218 or lesser degree. Sometimes, as in resonant interactions of wave-like motions, energy can move
 219 from one space- and time-scale to radically different ones. In conventional turbulence, interac-
 220 tions take place only between nearest-neighbor wavenumber scales, leading to the discussion of
 221 power laws, of which the Kolmogoroff $k^{-5/3}$ law is the best known, and where energy “cascades”
 222 from large to smallest scales. But in the ocean, some processes produce cascades in the oppo-
 223 site direction with differing power laws. Wave-like motions are often invoked statistically in a
 224 “wave-turbulence” approach (e.g., Polzin and Lvov, 2011).

225 Generalizing, $K_{1j}, K_{2j}, K_{j1}, K_{j2}$ are dominated by the balanced eddies in the open ocean,
 226 with the submesoscale becoming most important where isopycnal slopes are steep. K_{33}, K_{3j}, K_{j3}
 227 are usually controlled by breaking internal waves and provides the ultimate link between the
 228 stirring-dominated isopycnal terms and the ultimate dissipation. Generalizations about regions
 229 near lateral boundaries are not justified at the present time.

230 Power integrals

231 The most direct connection that can be made across the scales is that based upon “power
 232 integrals” as introduced by Stern (1975) and discussed by Joyce (1980). In a turbulent flow
 233 field, any structures in the C field would be dissipated at a rate,

$$\chi_C = 2\kappa_C \left(\left(\overline{\frac{\partial^2 C_{lm}}{\partial x_1^2}} \right)^2 + \left(\overline{\frac{\partial^2 C_{lm}}{\partial x_2^2}} \right)^2 + \left(\overline{\frac{\partial^2 C_{lm}}{\partial x_3^2}} \right)^2 \right), \quad (11) \quad \{\text{chi}\}$$

234 (the derivatives in Eq. (11) are negligible for scales larger than l_m) where the overbar denotes a
 235 space and/or time-average. Then in a statistically steady-state ocean, the input of this variance,
 236 integrated over the ocean is,

$$\iint_{\text{Surfaces}} C Q_C dA = \iiint_{\text{Vol.}} \chi_C dV, \quad (12) \quad \{\text{diss1}\}$$

237 and the integrals are taken over the ocean boundary area, particularly the surface, but including
 238 all generation boundaries, and the volume respectively, so that the generation-rate balances the
 239 dissipation rate. Q_C is a measure of the transfer of C to the ocean volume from the atmosphere
 240 and other boundaries under the assumption, in the steady-state, that

$$\iint Q_C dA = 0. \quad (13)$$

241 If Q_C and C are covarying, the resulting variance generated must be dissipated at the molecular
 242 scale, all physics of scale transformations being integrated out in Eq. (12). Generated variances
 243 on the left must be consumed in the term on the right at the molecular level, otherwise, the
 244 spatial variance of C would grow without bound. (See Schanze and Schmitt (2013) for the
 245 complications introduced by various nonlinearities.) The special case of $C = T$ corresponds to
 246 entropy conservation. Much of the dissipation, like the generation, occurs within surface the
 247 mixed-layer, but significant structures remain in the oceanic interior that must be removed at
 248 the molecular scale.

249 These arguments can be generalized to certain types of oceanic regional subvolumes. Using
 250 Eqs. (11), (12) to solve for κ_C gives values (for temperature and salinity) roughly consistent
 251 with those obtained from completely different methods (and with very different error bars). All
 252 observations do, however, remain highly uncertain; see Schneider and Bhatt (2000).

253 Consider now the analogue of Eq. (11) for the velocity field, defined as $\mathbf{v} = (u_1, u_2, u_3) =$
 254 (u, v, w) , and the latter being the Cartesian special case, but the first form permitting the use
 255 of indices. The dissipation of kinetic energy in the ocean, ultimately at similar molecular scales,
 256 is usually written as

$$\varepsilon = \frac{\nu}{2} \sum_{i,j=1}^3 \sum_{i',j'=1}^3 \left(\frac{\partial u_i}{\partial x_j} \frac{\partial u_{i'}}{\partial x_{j'}} \right) \quad (14)$$

257 ν is the ordinary molecular kinematic viscosity and which would again be dominated by the l_m
 258 spatial scale. If the turbulence lies mainly in isotropic motions at the smallest scales, the above
 259 simplifies to,

$$\varepsilon = \frac{15}{2} \nu \sum_{ij} \left(\frac{\partial u_i}{\partial x_j} \right)^2, \quad (15)$$

260 for any i, j . The connection between ε, χ , is exploited in the measurements described below.
 261 The ‘‘Kolmogoroff length-scale’’, $l_K = (\nu^3/\varepsilon)^{1/4}$ is a measure of the smallest turbulence scale

262 present in the fluid, and complete measurement of dissipation in the ocean would have to reach
 263 to distances of order 1 cm and less.

264 Large-scale forces acting on the ocean (e.g., winds and tides with time-scales ranging from the
 265 mean to short-lived squalls) are communicated to the ocean through processes which generate
 266 motions on all of the observed flow-scales. The wind and other generation processes can also
 267 destroy previously-produced structures, and thus Q must be interpreted as the *net* production
 268 in the boundary regions.

269 Vertical Mixing

270 *Quasi-Analytic Model Estimates*

271 The most influential attempt at estimating an element of \mathbf{K} was for the vertical (diapycnal)
 272 component K_{33} , with the advection-diffusion Eq. (9) written as the special case of a one-
 273 dimensional steady-state balance,

$$w \frac{\partial C}{\partial x_3} - \frac{\partial}{\partial x_3} \left(K_{33}(x_3) \frac{\partial C}{\partial x_3} \right) = 0. \quad (16) \quad \{\text{abyssa11}\}$$

274 For constant w , and K_{33} , the solution is a simple constant plus exponential,

$$C = C_0 + C_1 \exp(wx_3/K_{33}). \quad (17) \quad \{\text{abyssa12}\}$$

275 In widely influential paper, Munk (1966), noting the generally horizontal nature of isopycnals
 276 in the open Pacific Ocean, inferred w/K_{33} from fitting the Eq. (17) profiles of temperature and
 277 salinity below 1000 meters. Earlier, Wyrтки (1962) used passive oxygen and nutrient distributions
 278 which, however, involve interior sources and sinks having significant uncertainties. By also
 279 employing radiocarbon as a tracer with a decay term on the right hand side, Munk was able
 280 to separately distinguish $w \approx 10^{-7}m/s, K_{33} \approx 10^{-4}m^2/s$, assuming the parameters applied
 281 identically to temperature, salinity (or linearized density), and radiocarbon—as would be true
 282 in conventional turbulence. In the then convention of centimeter-gram-second (CGS) units, the
 283 result was the memorable (“canonical”) value $K_{33} = 1 \text{ cm}^2/s$.

284 Earlier “thermocline models” (e.g., Stommel and Webster, 1962; Welander, 1971) implicitly
 285 included Eq. (16) as a special case. These remarkable nonlinear solutions showed however, that
 286 “perfect fluid” (adiabatic, $\mathbf{K} = \mathbf{0}$) models gave a gross structure in the thermocline (roughly 300-
 287 1000m depth) similar to the observed temperatures and salinities and visually indistinguishable
 288 from those with significant vertical mixing. Munk’s “abyssal recipe” was widely mis-applied
 289 in that upper ocean region, where he had made no claims about K_{33} , and there arose a sense
 290 that a contradiction existed between the one-dimensional balance and a possible near-adiabatic

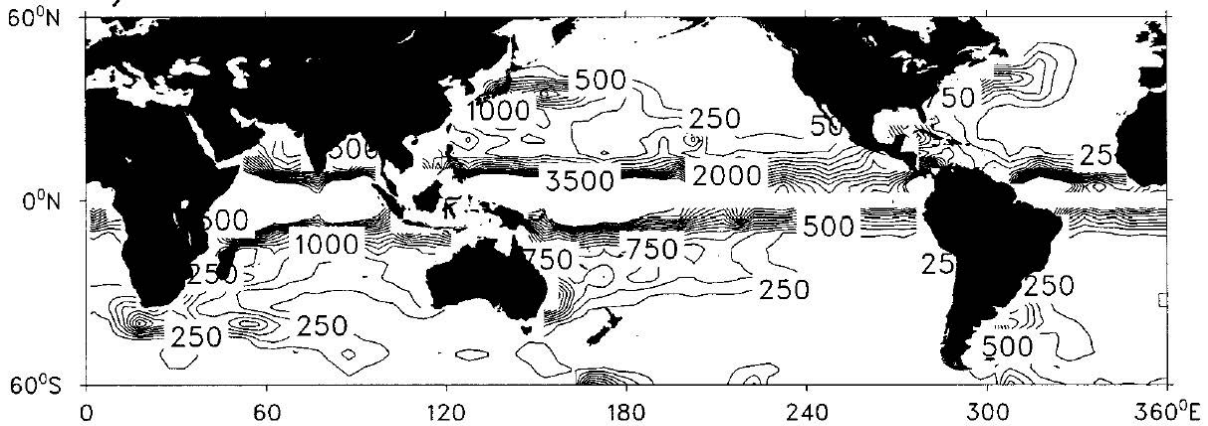


Figure 4: Estimated values of $K_{11} = K_{22}$ (m^2/s , Stammer, 1998) from altimetric variability. The equatorial region is omitted because of the breakdown of geostrophic balance there.

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291 upper ocean. The seeming-paradox was only resolved with the development of direct field
 292 measurements, taken up below.

293 Lateral Mixing

294 Lateral analogues in the horizontal dimension, for determining $K_{11,22}$, using large-scale observed
 295 horizontal gradients have a long history. An example is discussed by Needler and Heath (1975)
 296 and Hogg (1987) who, using the conspicuous feature of the Mediterranean Salt Tongue, did a
 297 simultaneous least-squares fit for both vertical and horizontal values. The latter found $K_{11} =$
 298 $K_{22} \approx 500 \text{ m}^2/\text{s}$, $K_{33} \approx 5 \times 10^{-5} \text{ m}^2/\text{s}$.

299 In more recent years, with the availability of near-global altimetric measurements of surface
 300 geostrophic flows, a number of attempts have been made to estimate $K_{11,22}$ directly from the
 301 eddy variability. Starting with Holloway (1986) and Stammer (1997), estimates range from
 302 about $250 \text{ m}^2/\text{s}$ in the quiescent eastern parts of the ocean to nearly $3500 \text{ m}^2/\text{s}$ in the eddy-
 303 intensified areas of the low-latitude western-boundary regions (Fig. 4). Numerous regional
 304 estimates have been published subsequently, with a particular emphasis in recent years on the
 305 eddy-rich Southern Ocean.

306 The lateral terms K_{1*} , K_{2*} where the $*$ denotes any of the 3-indices 1, 2, 3 are commonly
 307 identified primarily with the powerful kinetic energy of the balanced eddies, and within a mod-
 308 elling context have been the subject of a wide variety of proposals and tests. Recall, however,
 309 that the quasi-time average flow field has numerous spatial scales far smaller than the largest
 310 scale defining the general circulation. Boundary currents such as the Kuroshio and Gulf Stream
 311 on the west, the Equatorial Undercurrents and their associated shear fields, as well as flows on

312 eastern boundaries, are all capable of stretching and greatly extending the mixing volumes of
 313 tracers. But because the modelling community has focussed on parameterizing the mesoscale
 314 elements of \mathbf{K} , comparatively little is known about the mixing processes in strongly sheared
 315 boundary currents. More generally, the presence of strong horizontal flows (boundary currents,
 316 the circumpolar current) implies that the interaction of the balanced eddy field with those
 317 motions cannot be ignored. Note in particular (Ferrari and Nikurashin, 2010; Klocker et al.,
 318 2012) that the horizontal elements of \mathbf{K} are apparently suppressed near-surface at the core of
 319 the Antarctic Circumpolar Current, and presumably in other strongly sheared currents. In the
 320 same currents, excess mixing can occur at depth in the “critical levels” where the phase velocity
 321 of the eddy disturbance coincides with the velocity of the mean shear (Abernathey et al., 2010).

322 Lateral mixing has also been studied via the separation (dispersion) of clusters of pseudo-
 323 Lagrangian floats (Lacasse, 2008) as well as from data arising from the purposeful dye experi-
 324 ments. As with K_{33} , realistic numerical models of the ocean must accommodate a complex and
 325 highly detailed spatial structure in the lateral components of the mixing tensor. Whether the
 326 temporal variations, owing to variations in time of energy inputs to the ocean are significant
 327 and if so, on what time-scales, remains largely unknown.

328 **Vertical Mixing: Direct Field Measurements**

329 The discovery of what is now called “microstructure” on scales of centimeters in vertical profiles
 330 of temperature and salinity in the middle 1960s, through the newly available profiling instru-
 331 ments (see Baker, 1981) led to the realization that the observed structures were an indicator
 332 of the direct actions of molecular processes in the oceanic circulation. (“Fine structure” is
 333 commonly defined as appearing on the larger, meter, scales.)

334 Profiling instruments that measure the turbulent velocities have led to numerous direct
 335 estimates of diapycnal or vertical mixing rates. Osborn (1980) proposed that the relationship
 336 between the turbulent mixing coefficient and the measured dissipation could be calculated as,

$$K_{33} = \frac{\Gamma \langle \varepsilon \rangle}{N^2}, \quad (18) \quad \{\text{osborn}\}$$

337 where ε is defined in Eq. (12), brackets indicate an average, and $N = \sqrt{-(g/\rho) \partial\rho/\partial x_3}$ is the
 338 buoyancy frequency. Γ is the so-called mixing efficiency which is widely taken to be approxi-
 339 mately 0.2, but a growing body of evidence suggests that it can vary considerably (Inoue and
 340 Smyth, 2009; Lozovatsky and Fernando, 2013). Its physical interpretation is as the fraction of
 341 the turbulent energy that goes into mixing, noting that it must vanish in an unstratified fluid.
 342 Variations in Γ by factors of two or three are not usually the limiting uncertainty in calculations
 343 involving Eq. (18), but eventually it must not be treated as a universal constant.

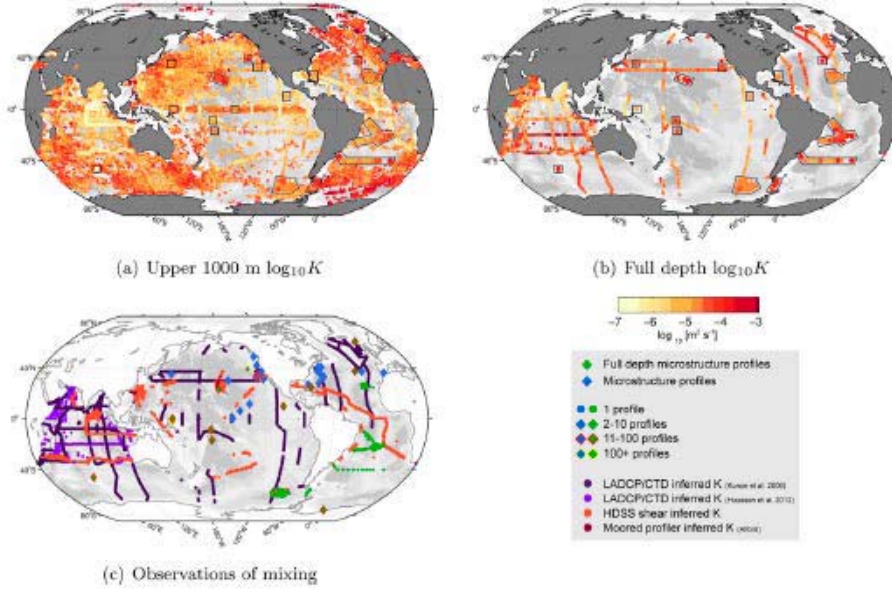


Figure 5: From Waterhouse et al. (2014) showing estimated values of K_{33} over the top 1000m (a) and averaged over the full depth (b) with many fewer measurements). (c) Shows the different types of measurements used for calculations. LADCP is lowered doppler current profiler; CTD is conductivity, temperature, depth; HDSS is hydrographic doppler sonar system.

{waterhouse_et}

344 An alternative measure (Osborn and Cox, 1972) is based directly on measurements of χ_T in
 345 the form,

$$K_{33} = \frac{\langle \chi_T \rangle}{2 \left(\frac{\partial \bar{T}}{\partial x_3} \right)}. \quad (19) \quad \text{{osborn-cox}}$$

346 The two methods should produce identical results where only temperature controls the stratifi-
 347 cation. Merrifield et al. (2016) discuss the differences when salinity becomes a factor. A major
 348 issue with Eqs. (18, 19) is that few instruments are capable of measuring the full spectrum
 349 of motions down to the Kolmogoroff scale, l_K in ε , or the even smaller equivalent Batchelor-
 350 scale, $l_B = \sqrt{\kappa^2 \nu / \varepsilon}$, for temperature. Thus the unresolved contributions to ε or χ_T must come
 351 from extrapolation. Some of these extrapolation formulae are quite complex. Waterhouse et
 352 al. (2014) have compiled a global calculation of K_{33} from a very large variety of velocity profile
 353 measurements and extrapolations, taken over the years, and a fixed value of Γ in Eq. (18), and
 354 are shown in Fig. 5). Upper ocean values are generally of order $10^{-5} \text{m}^2/\text{s}$, with higher values in
 355 the very high latitudes. Values at depth derive from very sparse sampling and tend to be higher
 356 over topographic features, e.g., the Mid-Atlantic Ridge in the South Atlantic Ocean.

357 Fig. 6 shows an approximate breakdown of their results by topographic roughness type.
 358 Abyssal plain values are nearly uniform with depth, while those over rough topography, and

359 particularly over the mid-ocean ridges show a pronounced increase towards the bottom. Any
360 calculation today of global mean abyssal values from these measurements would be highly un-
361 certain.

362 **Vertical Mixing: Volumetric Inverse Methods**

363 Profiling instruments produce measurements at single points, and even in Fig. 5 they remain
364 sparse in both space and time, conspicuously so at depth.. A different approach to deter-
365 mination of K_{33} comes from integrating geostrophic flows over large volumes lying between
366 isopycnals—and including boundary effects—in geostrophic box models. A closely related ap-
367 proach is through the β -spiral (e.g., Olbers et al., 1985; Fukumori, 1991).

368 A particularly simple example is that of Hogg et al. (1982) who found that the deep Brazil
369 Basin had to have a value of K_{33} of about $10^{-4}\text{m}^2/\text{s}$, not inconsistent with the abyssal recipes
370 canonical value. Later profiling measurements—Toole et al. (1997)—show the volumetric av-
371 erage there as being dominated by much higher values over and near the Mid-Atlantic Ridge
372 and lower values over the abyssal plain. Various box inversions (e.g., Ganachaud 2003, Lump-
373 kin and Speer, 2007, Macdonald et al., 2009) have produced basin-scale estimates ranging from
374 about 2×10^{-5} to 3×10^{-4} , depending upon the location and depth as well as assumptions
375 about detailed rates of flow. These results combine the effects of interior and lateral boundary/-
376 topography mixing rates. All have significant stated uncertainties. Many places (Ganachaud,
377 2003) show an increasing K_{zz} towards the bottom. A number of studies (e.g., Sloyan and
378 Rintoul, 2000) have estimated the net diapycnal flux.

379 Of particular interest have been the experiments with purposeful dye releases in the open
380 ocean (e.g., Watson et al., 2013) which tend to give higher mixing rates than have been inferred
381 from local profiling measurements—again probably best interpreted as an indication of higher
382 intensity mixing near boundary topographies of various sorts (Mashayek et al., 2016).

383 **Boundary Mixing**

384 Munk (1966) had noted in passing that oceanic mixing could be greatly enhanced at the sidewall
385 boundaries of the ocean. Stimulated by the intense motions generated on sloping boundaries by
386 the internal wave field, Phillips (1970) and Wunsch (1970) examined the flows at the boundaries
387 where no flux into the solid is permitted. Later examinations include that of Dell and Pratt
388 (2015). No simple separation into vertical and horizontal effects is possible in those regions.

389 The peculiar properties of internal waves on sloping boundaries, leading to intense break-
390 ing, and boundary layer instabilities (Ivey et al., 2000) seen in the laboratory led to fine and

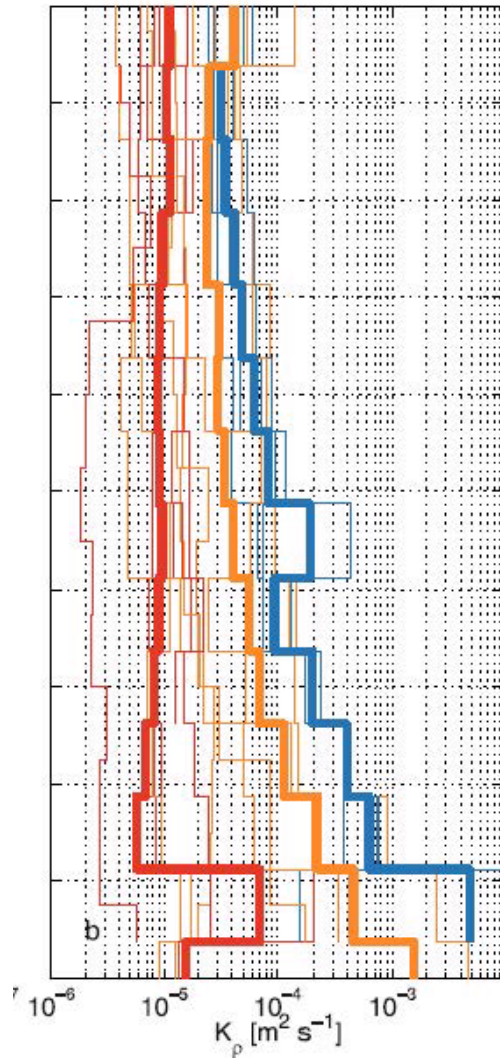


Figure 6: Diffusivity K_{33} for density as estimated by Waterhouse et al. (2013). Thick red curve is the average profile over smooth topography, thick orange curve is from rough topography and thick blue from ocean ridges. General increase towards the bottom is evident, particularly over the latter two regions. Thin lines show the scatter of each type.

{waterhouse_et

391 microstructure measurements, first near islands and seamounts (Wunsch, 1972; Hogg et al.,
392 1978; Eriksen, 1998) and to work on continental margins (e.g., Lamb, 2014). Direct experiments
393 with injected artificial dyes (e.g., Ledwell et al., 2016) have confirmed that excess mixing does
394 occur through various processes at continental boundaries. The collection of papers found in
395 Imberger (1998) provides a useful overview. Oceanic canyons, both on continental margins and
396 in the mid-ocean ridges also show strongly increased mixing, likely associated with internal wave
397 instabilities of various kinds (see Kunze et al., 2002, Thurnherr et al., 2005). Armi (1978) cal-
398 culated the effect of ambient flow and associated turbulence on sloping topography, but which
399 gave rise to debate (see Garrett et al., 1993) concerning the influence of that mixing on the
400 interior stratification. At the present time, it does appear that sloping boundary turbulence
401 and consequent mixing are dominated by the complex motions and breaking from the internal
402 wave field, including the internal tide. Whether the maximum effective mixing rates occur in the
403 boundary layers, or in the interior region above the boundary layer is not so clear. Garrett et al.
404 (1993) have argued that boundary layers tend to mix already mixed fluid and so are relatively
405 ineffective compared to the enhanced breaking of internal waves just above the bottom.

406 A sweeping generalization would be that the Munk canonical value for a vertical mixing
407 coefficient of about $10^{-4}\text{m}^2/\text{s}$ applies primarily to the deep ocean (somewhere deeper than the 1
408 km depth he focussed on), as an average of much smaller interior values (below $10^{-5}\text{m}^2/\text{s}$) and
409 much higher values over topographic features, with generally increasing values as the topography
410 (on many scales) is approached from above (See Toole et al., 1994; Munk and Wunsch, 1998).
411 Regions with very small values of inferred K_{33} imply correspondingly very small values of the
412 vertical velocity, w , in those places (see the next section), with major consequences for the
413 behavior of the large-scale circulation. Furthermore, the strong increase of K_{33} toward the
414 oceanic floor, away from the abyssal plains, gives rise to somewhat unexpected overall behavior
415 through terms such as $\partial K_{33}/\partial x_3$ (Ferrari et al., 2016) which act in the advection-diffusion
416 equation as an equivalent abyssal vertical velocity.

417 Boundary mixing is included in the inverse box model inversions where hydrographic lines
418 terminate in topographic features, but not in most of the profiling measurements. Comparisons
419 of volume averages from box models with point measurements from profilers are thus difficult.

420 In the region above 1-2 km in the water column, the dominant values of $K_{33} \leq 10^{-5}\text{m}^2/\text{s}$
421 are consistent with the thermocline theories treating the motion as nearly non-diffusive. Much
422 recent attention has thus focussed on the outcrop regions of the isopycnals/neutral surfaces in the
423 Southern Ocean, where direct wind mixing, within the surface mixed-layer/Ekman layer, likely
424 provides most of the transformation of water properties required by the observed upper-ocean
425 circulation.

426 Exchanges between the oceanic interior and fluid in the boundary layers of the sidewalls
 427 (recalling that their slope is usually less than 3°) requires special treatment in its own right
 428 (e.g., Phillips et al., 1986; Garrett et al., 2013; Dell and Pratt, 2015) and may (Ferrari et al.,
 429 2016) control the interior circulation. Attention in the next several years is likely to further
 430 focus on the interaction of ocean flows on all scales with boundary features of all types.

431 **Consequences of Spatial Variations**

432 The existence of regions of strongly enhanced values of K_{ij} , particularly the orders of magnitude
 433 found in K_{33} , greatly complicates quantitative calculations of effective rates of mixing. Not only
 434 does the spatial variation have important consequences for the ocean circulation on the largest,
 435 L , scale, but the residence time of fluid within those “hotspots” is a function of the large-scale
 436 circulation itself (Mashayek et al., 2016).

437 The simplest dependence on the spatial structure can be seen from the important dynamical
 438 relationship in a mainly geostrophically balanced ocean in the linear vorticity equation,

$$\beta v = f \frac{\partial w}{\partial z} \quad (20)$$

439 where v is the meridional velocity, and β is the local meridional derivative of the Coriolis pa-
 440 rameter, f . To have a meridional flow requires a finite w (the conventional Eulerian vertical
 441 component) with finite derivative. On the other hand, in regions of nearly horizontal isopycnals,
 442 Eq. (16) shows that the existence of w depends directly upon the existence of a finite K_{33} . Weak
 443 values of vertical mixing, in the absence of steeply sloping isopycnals, imply weak or non-existent
 444 w , and hence a strong limit on any meridional flow.

445 In the “abyssal recipe” of Eq. (16), a steady-state e.g., of temperature, $C = T$ is maintained
 446 with a constant K_{33} as a balance between the downward diffusion of heat from the high temper-
 447 atures near the surface, and the upward advection cold water from below. But if the assumption
 448 of a vertically constant K_{33} is abandoned, the situation can be very different. Suppose by way
 449 of example that the abyssal ocean has a constant temperature gradient, $\partial T / \partial x_3 = T'_0$, then Eq.
 450 (16) becomes

$$w = \frac{\partial K_{33}}{\partial x_3} \quad (21) \quad \{\text{wfromk}\}$$

451 and if, as now seems clear (Fig. 5), K_{33} is often largest near the ocean bottom, $w < 0$ and
 452 the balance is now that of the *downward advection* of temperature against the *upward diffusion*
 453 of cold water.² An implication is the reversal of the standard picture of an upward movement

²The argument that true mixing can only increase the potential energy applies only to the net effects over the entire volume.

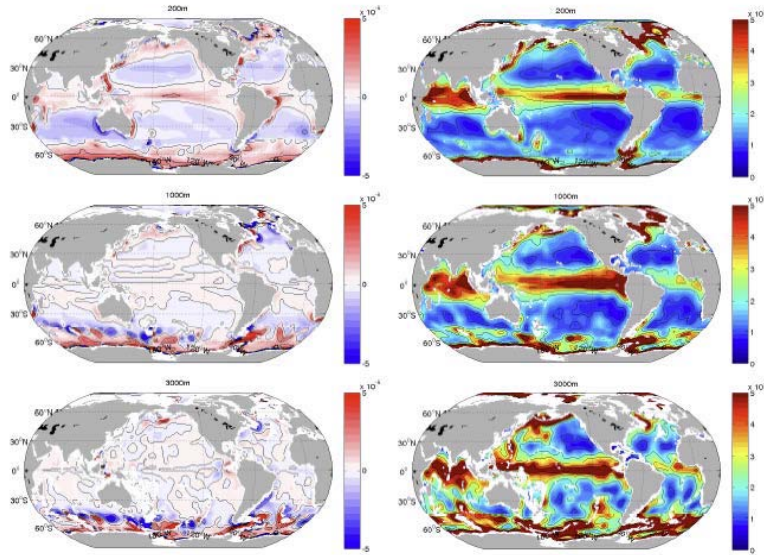


Figure 7: A 20-year average estimated Eulerian w from a highly constrained general circulation model (a “state-estimate”; see Forget et al., 2015). Values at three depths are shown in the left-column, temporal standard deviations in the right-column. 200m values are dominated by the Ekman pumping of the mean wind, while the deepest values at 3000m are topographically controlled. Intermediate depth displays a combination of both fields. (X. Liang, 2015, personal communication). The complex reversals show that some universally valid description about the nature of vertical ocean fluxes is likely not possible.

{liang_vertica

454 of water balancing high latitude convection of dense, deep water to the abyss. The product of
 455 the density profile with that of K_{33} determines the sign of w . A more general derivation of this
 456 result, with a discussion of the potential consequences for the general circulation, is given by
 457 Ferrari et al. (2016).

458 The implied values of w from Eq. (21) will have a lateral variability that is as large or
 459 larger than that seen in K_{33} itself. Fig. 7 displays a 20-year average and its standard deviation
 460 obtained from a state estimate (an ocean general circulation model fit to the data by constrained
 461 least-squares). Among other implications is that the Eulerian w field in the ocean is expected
 462 to be extremely noisy; Fig. 7 displays an estimate of w at three depths. The spatial structures,
 463 particularly at depth, are complex even with 20-years of averaging. Liang et al. (2016) showed
 464 equivalent maps of the net time-averaged $w^* = w + w_{eddy}$ arising both from the Eulerian and
 465 eddy-induced values. No global generalization, even about just the open ocean values of w or
 466 w^* , is likely to be accurate. No reconciliation yet exists of the inferences from the Munk (1966)
 467 model, and those deriving from the intense spatial variability.

468 **Upper Ocean/Mixed Layer**

469 The surface layers of the ocean, in direct contact with the atmosphere have a specialized mixing
470 literature of their own (D’Asaro, 2014). Surface outcrop areas of isopycnals visible in Fig. 3
471 and in other oceans are of particular importance. Exchanges between the mixed layer and the
472 ocean below have focussed on the influence of the sub-mesoscale, particularly as it influences
473 the nutrient transfers, and hence the biology (e.g., Mahadevan, 2016).

474 **Vertical Mixing: The Energy Problem**

475 In the underlying Navier-Stokes equations such as (1), κ_C derives from the vibrational and other
476 motions of molecules powered by the internal energy of the fluid. Consider that an isolated
477 thermally stratified fluid will, after sufficiently long time, become uniform in temperature, with
478 a consequent increase in its potential energy—the center of mass having been moved upward and
479 with the internal energy correspondingly reduced. In numerical models, where the coefficients
480 such as \mathbf{K} are commonly introduced as ad hoc parameters representing the action of eddies on
481 the larger scale L , no explicit source of energy has been specified. Beginning in the late 1990s,
482 attention began to turn toward understanding of the complete oceanic energy budget including,
483 specifically, the source of energy powering the stirring carrying the mixing the final step toward
484 the molecular scale.

485 Energy suppliers for the turbulent stirring and mixing are identical to those powering the
486 overall general circulation and thus a clean separation of sources is not possible. Such overall
487 power sources include: (1) the wind; (2) tides; (3) precipitation and evaporation; (4) heat
488 exchange with the atmosphere; (5) atmospheric pressure changes. Attaching numbers with
489 useful accuracy to these various processes has proved challenging, and the reader is referred for
490 more detail to the article on ocean energetics. A summary would be that (3) and (5) are minor
491 contributors, (1) and (2) are very important, and that evaluating (4) has proven problematic—as
492 it appears to depend strongly on the flows established by (1) as well. Input of energy from the
493 wind is thought to take place primarily on the largest scales, and on the scale of oceanic inertial
494 motions (near the Coriolis frequency, f). Much of the energy input appears first as an increase in
495 oceanic potential energy, prior to its release into kinetic energy by a variety of pathways. Winds
496 are believed to be the major power source for the balanced eddies via baroclinic instabilities, but
497 then can act to destroy them. Tidal input of energy is believed to occur primarily through the
498 conversion at topography of the ordinary surface tides into internal tides on scales far smaller
499 than L .

500 Understanding the energetics of mixing then depends upon determination of the pathways

501 by which the energy inputs carry energy down to the point where molecular processes can begin
502 to work efficiently. These pathways and their importance and locales remain poorly quantified.
503 Attention has focussed on internal waves—both interior breaking (see e.g., Munk, 1981; Thorpe,
504 2010), and the very intense motions they induce on and over slopes of all kinds.

505 Most of the kinetic energy of the ocean circulation lies in the balanced eddies and they
506 are an obvious source for conversion into smaller scales. As compared to the internal wave
507 field, their vertical and horizontal scales tend to be larger, but in topographic interactions they
508 are capable of generating a large variety of much smaller scales including internal waves and
509 boundary layers. The lifetime of the eddy field has been variously estimated as roughly 6
510 months (Ferrari and Wunsch, 2009), but where and how the corresponding dissipation occurs
511 is not fully quantified (Zhai et al., 2010; Clément et al., 2016). The direct coupling between
512 geostrophically balanced motions and internal waves is weak (Vanneste, 2013). Also a strong
513 tendency exists in rotationally dominated turbulence to drive energy “upscale”, that is to larger,
514 rather than smaller spatial scales (e.g., Vallis, 2006). But turbulent bottom friction, topographic
515 interactions, flow separation at bottom features, all remain viable dissipation candidates.

516 **General Circulation Models (GCMs)**

517 Oceanic kinetic energy (what moves tracers around) is dominated by the balanced eddies. As
518 modelling power has grown, the greatest efforts have been directed at the detailed numerical
519 representation of the balanced eddies in models. In particular, many global models fail to resolve
520 the eddy field, and the numerical parameterization of the effects of balanced eddies has generated
521 a voluminous literature (see the collection of papers in Hecht and Hasumi, 2007). In numerical
522 models the mixing tensor (Eq. 10) is generally formulated in isopycnal coordinates, reflecting the
523 anisotropy of diapycnal and isopycnal mixing. In so-called z -coordinate models the resulting
524 tensor is then transformed approximately back to geodetic coordinates. The re-transformation
525 is not required for models written in isopycnal coordinates.

526 Because lateral motions are much more intense than vertical ones, numerical difficulties arise
527 in preventing artificial vertical diffusivities arising from the horizontal motions, distinguishing
528 between those fields (temperature and salinity) which determine the isopycnals, and the passive
529 tracers that make no such contribution. This complex subject is covered in the textbooks by
530 Haidvogel and Beckmann (1999); Griffies (2004), Vallis (2006), Olbers et al. (2012).

531 The largest literature focuses on the best numerical representation of \mathbf{K} —in all three-
532 dimensions—and its equivalents (the problem of “sub-grid scale parameterization”). In the
533 process, the distinction between stirring and mixing has largely been lost, with the molecular
534 scales assumed to take care of themselves—between the actions of \mathbf{K} , and with the inescapable

535 numerical dissipation becoming a surrogate for the true molecular mixing. Much (not all) of
536 the eddy effect is represented in terms of an equivalent lateral velocity (sometimes called the
537 bolus velocity), and is calculated from the anti-symmetric components of tensor \mathbf{K} (see espe-
538 cially Flierl and McGillicuddy, 2002, Section 7; the “skew-flux”), and is related closely to the
539 transformed Eulerian mean (TEM). A common numerical method relies upon extensions and
540 modifications of the proposal of Gent and McWilliams (1990). Because the method, and most
541 of its variants, is focussed on baroclinic instability processes, it removes potential energy from
542 the fluid by flattening isopycnals, rather than increasing it as in mixing. It is thus a scheme for
543 spreading, and not mixing. An explicit dissipation picture for the energy released is lacking.

544 Many of the isopycnals in the upper 1000m of the ocean—above the domain of Munk’s
545 “abyssal recipes”—outcrop at the sea surface in the Southern Ocean (Fig. 3). There the intense
546 turbulence existing in the surface boundary layer, derived directly from the locally powerful
547 wind field, can provide an easy route for mixing the fluid across these density surfaces. (see
548 e.g., the modelling study of von Storch et al., 2007). In the abyss, where the isopycnals do not
549 outcrop to the surface, direct wind-forcing does not act and diapycnal mixing is dependent upon
550 the wave breaking mechanisms.

551 In a hypothetical ocean model that was capable of directly resolving the centimeter and
552 smaller scales, l , important in $\kappa_C \nabla^2 C$, no parameterizations would be required. Taking the
553 requisite length arbitrarily as 1 mm, vertically and horizontally, a global model would require
554 of order 10^{27} grid points or Fourier components, and thus the need for statistical mechanics
555 and parameterizations will exist indefinitely. Global models in which the balanced eddy field
556 is adequately resolved are on the horizon, but the mixing scales, l_m , remain far beyond direct
557 reach.

558 **Summary Discussion**

559 The remarkable spatial variations now known to exist in ocean mixing processes have many
560 implications for understanding and potential prediction of the ocean circulation and all of its
561 climate consequences. With the coupling amongst motions at all time and space scales, discus-
562 sion of mixing can hardly be done without a long excursion into all ocean flows and energetics.
563 Spatial variations in effective mixing have powerful implications for the general circulation it-
564 self (defined as a time average over several decades) implying a very different picture than was
565 accepted until quite recently.

566 In seeking gross generalizations, it appears that the tide—directly generating internal waves—
567 and the wind generated instabilities and inertial motions, produce the turbulence that mixes the
568 ocean. Roughly speaking, the equivalent vertical or diapycnal or dianeutral mixing coefficients

569 range from comparatively low average values in the upper 1000m of about $10^{-5}\text{m}^2/\text{s}$, to an order
570 of magnitude larger below that depth. Very much higher values appear in the surface mixed layer
571 with distinct physics. Values increase in the abyss, over certain types of topographic features,
572 by two or more additional orders of magnitude. The implications are that much of the upper
573 ocean can be treated as a near-perfect fluid, with the necessary mixing occurring at the sea
574 surface at high latitudes. In contrast, the very intense mixing over topographic features implies
575 that vertical exchanges and hence the meridional oceanic motions are confined to very special
576 regions, including lateral boundaries of all types. A general increase in diapycnal mixing toward
577 the sea floor in the open ocean implies a net downward flux of temperature generally, with
578 part of the upwards return flow being a lateral boundary phenomenon—and which represents a
579 radical change in understanding of the deep circulation.

580 Determination of the influence of the ocean on future climate change requires quantitative
581 prediction of the ways in which energy exchanges with the atmosphere will shift, and in turn
582 how those translate into modifications of the general circulation, its storage and transfers of
583 heat, freshwater, carbon, oxygen, etc. The gross ocean stratification will also change, along
584 with the mixing energetics and rates, as well as will the meteorological interactions. *A priori*
585 calculation of the overall stratification and resulting movement of climate properties in the ocean
586 past, present, and future remains a major challenge.

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